

An approach for a polychromatic generalized Lorenz-Mie theory

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Abstract

In this work a polychromatic version of the generalized Lorenz-Mie theory *stricto sensu* (GLMT) is derived. In this new formalism, arbitrary time-dependent fields are expanded into partial waves using Bromwich scalar potentials and, instead of the usual expansion coefficients – the beam shape coefficients (BSCs) – found in the monochromatic GLMT, now one finds *field shape spectra* (FSSs) which are intrinsically frequency-dependent. Expressions for the incident, scattered and internal fields are presented, and it is shown how physical quantities defined and expressed in the monochromatic GLMT in terms of the BSCs are modified and redefined in terms of the FSSs in polychromatic light scattering problems, like scattered intensities and phase angles, absorption, extinction and scattering cross-sections, and radiation pressure cross-sections.

Keywords: generalized Lorenz-Mie theory, light scattering, polychromatic light

1. Introduction

2 The experimental advent of the laser by Maiman, in 1960 [1], allowed
3 physicists and engineers to produce, in a controlled manner, transversely
4 localized coherent light, opening a new era of investigations in optics and
5 photonics. Concomitantly, it also forced theoreticians to devise, among oth-
6 ers, new formulations to incorporate such transversally-localized, Gaussian-

7 like light into existing light scattering “theories”, like the Lorenz-Mie theory
8 (LMT) [2, 3, 4].

9 The LMT deals with the scattering and absorption of monochromatic
10 plane waves by homogeneous, spherical particles of radius a of arbitrary
11 value relatively to the wavelength λ of the incoming light. In this framework,
12 the plane wave is expanded into a set of partial waves using spherical wave
13 functions. Scattered and internal fields, absorption, scattering and extinction
14 cross-sections and radiation pressure forces, among other physical quantities
15 of possible interest, follows this infinite-sum structure, although truncations
16 might become plausible depending on the ratio a/λ . Of course, the larger the
17 number of non-negligible terms in the expansion, the more computationally
18 demanding will be the calculations [5]. This number is determined by the
19 size of the scatterer and its electromagnetic parameter relative to the host
20 medium in which it is immersed, or might depend on the spatial domain of
21 reconstruction, e.g., in computing the fields in a certain region.

22 The simplest extension of the LMT would be the one that replaces the
23 non-physical plane waves by arbitrary-shaped beams like Gaussian-like beams
24 (fundamental TEM₀₀ mode) produced by laser sources [6, 7, 8, 9, 10, 11,
25 12]. Although light produced in this manner cannot rigorously be assumed
26 monochromatic, today’s technology allows us to speak of laser beams as
27 single-frequency, temporal coherent wave fields without introducing observ-
28 able theoretical deviations from practical light-matter problems. Instead of
29 talking about a very narrow spectrum centered at λ , the laser source is as-
30 sumed to emit light at that specific wavelength. This simplest extension of
31 the LMT is widely known as the generalized LMT (GLMT) *stricto sensu*
32 [13, 14, 15, 16], or simply GLMT in the context of the present work, see
33 Refs. [17, 18, 19, 20, 21, 22] for review, historical and scientific facts.

34 From a theoretical point of view, the fundamental difference between the
35 LMT and the GLMT lies on the expansion coefficients of the partial wave
36 expansion, which now must be altered in magnitude and phase in order to
37 accommodate for the new spatial field distribution [23, 24, 25, 26]. Such
38 expansion coefficients, known in the literature as the beam shape coeffi-
39 cients (BSCs), carry, therefore, all the spatial nuances of an arbitrary-shaped,
40 monochromatic beam, relative to a plane wave. Incidentally, the GLMT was
41 first applied to Gaussian beams, since these were the most common type of
42 laser sources available in optics laboratories during the 70’s and 80’s. It must
43 be clear, however, that since the BSCs can be modeled at will, and that, re-
44 gardless of their values, a true Maxwellian beam will always be produced,

45 they apply to any envisioned structured light, contrary to past claims now
46 firmly rebutted (see, for instance, the Introduction of Ref. [21]).

47 One of the most successful versions of the GLMT has its mathematical
48 roots on Bromwich scalar potentials (BSPs) rather than on Hertz or magnetic
49 vector potentials [15, 27, 28]. The BSPs can be defined for certain classes of
50 coordinate systems for which the coordinates are separable and whose scale
51 factors obey specific extra conditions. The spherical coordinate system meets
52 all such required conditions. Although BSPs do not satisfy a wave equation,
53 the electromagnetic fields can be easily extracted from them. In the GLMT,
54 time-harmonic fields are taken for granted, so that the formalism is derived
55 in terms of phasor fields starting from TM (Transverse Magnetic) and TE
56 (Transverse Electric) BSPs assuming a fixed but otherwise arbitrary angular
57 frequency ω .

58 Despite its recognized success over the past decades, the mathematical
59 framework of the GLMT is still limited to monochromatic arbitrary-shaped
60 beams. No systematic study is available showing the step-by-step procedure
61 to incorporate polychromatic electromagnetic fields into the analysis of scat-
62 tered intensities, phase angles, radiation pressure forces and torques, and so
63 on. A formalism for such a *polychromatic* GLMT would help in the anal-
64 ysis, for instance, of the interaction between spherical particles with RGB
65 (red, green, blue) laser system. Optical pulses with large bandwidth and
66 their interaction with small particles could also be readily investigated, in-
67 cluding femtosecond pulsed light generated by non-continuous wave lasers
68 [29, 30, 31, 32, 33, 34, 35]. Finally, such a framework would also pave
69 the way for extensions such as polychromatic Extended Boundary Condi-
70 tion Method (EBCM) – not to be confused with T-matrix method, insofar as
71 GLMT, among others, is as well a T-matrix method [18, 20], polychromatic
72 light-matter interactions with multilayered spheres, ellipsoidal and cylindri-
73 cal particles, aggregates of scatterers and so on.

74 This paper is therefore devoted to the task of presenting a systematic ap-
75 proach to a polychromatic GLMT. The derivation is based on BSPs (which
76 could be possibly translated to a more compact but less readable theory in
77 terms of Vector Spherical Wave Functions, e.g. Ref. [18] and “Supplemental
78 Material for the Second Edition” in Ref. [28]), and extended formulas for
79 the incident, scattered and internal fields are presented in Section 2. As will
80 be seen, the BSCs become, as they should, frequency-dependent, now called
81 *field shape spectra* (FSSs). In Section 3, new expressions for scattered inten-
82 sities and phase angles, extinction, scattering and absorption cross-sections,

83 and radiation pressure cross-sections are derived which can be applied to
 84 continuous or discrete frequency-dependent incident optical fields. Finally,
 85 Section 4 is a conclusion.

86 2. Generalities

87 Let us start by assuming some basic notation for the GLMT for monochro-
 88 matic arbitrary-shaped beams. In what follows, by GLMT we mean the
 89 GLMT *stricto sensu*, that is, valid for homogeneous spherical scatterers. A
 90 Cartesian coordinate system (x, y, z) is attached to a sphere of radius a ,
 91 which is centered at its origin O_P . Corresponding cylindrical and spherical
 92 coordinate systems (ρ, φ, z) and (r, θ, φ) are assumed to be related to (x, y, z)
 93 as usual. The scatterer has a relative refractive index of $M = [\mu_{sp}\epsilon_{sp}/\mu\epsilon]^{1/2}$,
 94 μ and ϵ (μ_{sp} and ϵ_{sp}) being, respectively, the permeability and permittivity
 95 of the host medium (particle).

96 In the GLMT based on Bromwich Scalar Potentials (BSPs), we assume
 97 classes of orthogonal curvilinear coordinate systems with coordinates x^k ($k =$
 98 $1, 2, 3$) and scale factors e_k , such that [28]:

$$e_1 = 1, \quad (1a)$$

$$\partial_1 \left(\frac{e_2}{e_3} \right) = 0, \quad (1b)$$

100 with $\partial_k \equiv \partial/\partial x^k$.

101 From Maxwell's equations in the time domain, BSPs $U(\mathbf{r}, t)$ must satisfy
 102 a differential equation of the form [28]:

$$\partial_1^2 U(\mathbf{r}, t) - \mu\epsilon \partial_t^2 U(\mathbf{r}, t) + \frac{1}{e_2 e_3} \left[\partial_2 \frac{e_3}{e_2} \partial_2 + \partial_3 \frac{e_2}{e_3} \partial_3 \right] U(\mathbf{r}, t) = 0. \quad (2)$$

103 For spherical coordinates, Eq. (2) can be specialized to read:

$$\partial_r^2 U(\mathbf{r}, t) - \mu\epsilon \partial_t^2 U(\mathbf{r}, t) + \frac{1}{r^2 \sin \theta} \partial_\theta [\sin \theta \partial_\theta U(\mathbf{r}, t)] + \frac{1}{r^2 \sin^2 \theta} \partial_\varphi^2 U(\mathbf{r}, t) = 0. \quad (3)$$

104 As is well known from the GLMT, Eq. (3) is not a wave equation for
 105 $U(\mathbf{r}, t)$. However, this partial differential equation do admit solutions of the
 106 form:

$$U(\mathbf{r}, t) = rR(kr)\Theta(\theta)\Phi(\varphi)T(t). \quad (4)$$

107 Notice the explicit appearance of time in the present approach, which
 108 contrasts with the one adopted in the traditional GLMT for monochromatic
 109 fields [15]. Since we are working explicitly in the time domain with arbitrary
 110 instantaneous fields, a temporal function $T(t)$ has been explicitly considered
 111 in the method of separation of variables.

112 Substituting Eq. (4) in (3), dividing the resulting expression by $U(\mathbf{r}, t)$
 113 and temporarily omitting the arguments, one then has:

$$\frac{1}{rR} d_r^2 (rR) - \mu\epsilon \frac{1}{T} d_t^2 T + \frac{1}{r^2 \sin \theta} \frac{1}{\Theta} d_\theta [\sin \theta d_\theta \Theta] + \frac{1}{r^2 \sin^2 \theta} \frac{1}{\Phi} d_\varphi^2 \Phi = 0. \quad (5)$$

114 In Eq. (5), the notation $d_k^i = d^i/dx^k$ has been introduced, in which x^k
 115 now stands for r , θ , φ and t . Isolating the term containing T ,

$$\frac{1}{rR} d_r^2 (rR) + \frac{1}{r^2 \sin \theta} \frac{1}{\Theta} d_\theta [\sin \theta d_\theta \Theta] + \frac{1}{r^2 \sin^2 \theta} \frac{1}{\Phi} d_\varphi^2 \Phi = \mu\epsilon \frac{1}{T} d_t^2 T. \quad (6)$$

116 The r.h.s. of Eq. (6) depends exclusively on t , while the l.h.s. depends
 117 on other independent spatial variables. Therefore, both sides must be equal
 118 to a possibly complex constant which we shall call $-k^2$, so that:

$$\mu\epsilon \frac{1}{T} d_t^2 T = -k^2 \quad (7)$$

119 After rearranging Eq. (7) and introducing the definition $k^2 \equiv k^2(\omega) =$
 120 $\omega^2 \mu\epsilon$:

$$d_t^2 T + \omega^2 T = 0. \quad (8)$$

121 As expected, the solutions to the temporal contribution in the method of
 122 separation of variables are simple harmonic functions $\exp(i\omega t)$, where ω is
 123 the angular frequency and, therefore, a continuous variable in general.

124 Having solved for $T(t)$, Eq. (5) is now multiplied by r^2 and rearranged so
 125 as to read as:

$$\frac{r}{R} d_r^2 (rR) + k^2 r^2 = -\frac{1}{\sin \theta} \frac{1}{\Theta} d_\theta [\sin \theta d_\theta \Theta] - \frac{1}{\sin^2 \theta} \frac{1}{\Phi} d_\varphi^2 \Phi = 0. \quad (9)$$

126 Except for the fact that we are not assuming time-harmonic functions
 127 with a particular frequency $\omega = \omega_0$, Eq. (9) is identical in form with Eq. (2.64)
 128 of Ref. [28], so that the solution to the above equation can be directly repro-
 129 duced from Eq. (2.91) in the aforementioned reference.

130 Letting $U(\mathbf{r}, \omega) = rR(kr)\Theta(\theta)\Phi(\varphi)$ [remember that $k \equiv k(\omega)$], one then
 131 has:

$$U(\mathbf{r}, \omega) = \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} \frac{c_{nm}}{k} \begin{pmatrix} \psi_n(kr) \\ \xi_n(kr) \end{pmatrix} P_n^{|m|}(\cos \theta) e^{im\varphi}. \quad (10)$$

132 In Eq. (10), $\psi_n(z) = zj_n(z)$ and $\xi_n(z) = zh_n^{(2)}(z)$ are Ricatti-Bessel func-
 133 tions, with $j_n(z)$ and $h_n^{(2)}(z)$ being, respectively, spherical Bessel functions of
 134 the first kind and spherical Hankel functions of the second kind (the latter
 135 functions may also be called spherical Bessel functions of the fourth kind, see
 136 Chapter 2 in Ref. [28]). The associated Legendre functions $P_n^{|m|}(\cos \theta)$ are
 137 defined in accordance with Hobson's notation [36]. Finally, the pre-factors
 138 c_{nm} depend on n , m and possibly on k or, equivalently, ω .

139 Therefore, the general form for time-dependent BSPs $U(\mathbf{r}, t)$ is given by:

$$\begin{aligned} U(\mathbf{r}, t) &= U(\mathbf{r}, \omega) e^{i\omega t} \\ &= \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} \frac{c_{nm}}{k} \begin{pmatrix} \psi_n(kr) \\ \xi_n(kr) \end{pmatrix} P_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t}, \end{aligned} \quad (11)$$

140 which is simply the usual BSP in the monochromatic GLMT multiplied by
 141 the time-harmonic factor $\exp(i\omega t)$.

142 At this point, one may define a *generalized* BSP by adding up all possible
 143 solutions to the partial differential equation, Eq. (6):

$$U(\mathbf{r}, t) = \sum_j \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} \frac{c_{nm}}{k} \begin{pmatrix} \psi_n(kr) \\ \xi_n(kr) \end{pmatrix} P_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t}. \quad (12)$$

144 In Eq. (12), the symbol \sum_j can actually represent either discrete or con-
 145 tinuous (integrals) sums, or compositions of both. In general, therefore, one
 146 could then propose:

$$\begin{aligned} U(\mathbf{r}, t) &= \sum_{j=1}^J \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} \frac{c_{nm}^j}{k_j} \begin{pmatrix} \psi_n(k_j r) \\ \xi_n(k_j r) \end{pmatrix} P_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega_j t} \\ &+ \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} \frac{c_{nm}(\omega)}{k(\omega)} \begin{pmatrix} \psi_n[k(\omega)r] \\ \xi_n[k(\omega)r] \end{pmatrix} P_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega, \end{aligned} \quad (13)$$

147 OR

$$U(\mathbf{r}, t) = \sum_{j=1}^J U_d(\mathbf{r}, \omega_j) e^{i\omega_j t} + \int_{-\infty}^{\infty} U_c(\mathbf{r}, \omega) e^{i\omega t} d\omega. \quad (14)$$

148 In Eq. (13), $c_{nm}^j \equiv c_{nm}(\omega_j)$ and $k_j \equiv k(\omega_j) = \omega_j \sqrt{\mu(\omega_j)\epsilon(\omega_j)} = 2\pi/\lambda_j$,
 149 with $\lambda_j \equiv \lambda(\omega_j)$ being the wavelength of the j -th monochromatic field with
 150 angular frequency ω_j . Notice that we have explicitly assumed that the con-
 151 stituent parameters of the host medium might change with (angular) fre-
 152 quency. The lower indices “ d ” and “ c ” stands for *discrete* and *continuous*
 153 BSPs $U_d(\mathbf{r}, \omega_j)$ and $U_c(\mathbf{r}, \omega)$, respectively, whose general forms can be taken
 154 from Eq. (13).

155 Equation (13), or (14), is the most general form for BSPs for light fields
 156 composed of both sums of monochromatic (single frequency) beams and poly-
 157 chromatic optical waves. However, since

$$\sum_{j=1}^J U_d(\mathbf{r}, \omega_j) e^{i\omega_j t} = \int_{-\infty}^{\infty} \left[\sum_{j=1}^J U_d(\mathbf{r}, \omega_j) \delta(\omega - \omega_j) \right] e^{i\omega t} d\omega, \quad (15)$$

158 where $\delta(z)$ is the Dirac delta function, the term under brackets in Eq. (15)
 159 can be merged with $U_c(\mathbf{r}, \omega)$ in Eq. (14), allowing us to define the BSP spectra
 160 $U(\mathbf{r}, \omega)$ as follows:

$$U(\mathbf{r}, t) = \int_{-\infty}^{\infty} U(\mathbf{r}, \omega) e^{i\omega t} d\omega. \quad (16)$$

161 As given by either Eq. (14) or (16), $U(\mathbf{r}, t)$ can be viewed as a *general-*
 162 *ized* BSP. The second term in the r.h.s. of Eq. (14) corresponds, except for
 163 an absent multiplicative factor of $1/2\pi$, to the inverse Fourier transform of
 164 $U_c(\mathbf{r}, \omega)$, that is, of the continuous part of $U(\mathbf{r}, t)$. Once the discrete part of
 165 the generalized BSP, $U_d(\mathbf{r}, \omega)$, is incorporated into a single spectra $U(\mathbf{r}, \omega)$,
 166 Eq. (16) is simply a Fourier transformation process of an arbitrary polychro-
 167 matic BSP, as expected. Notice that, in writing the r.h.s. of Eq. (15) we have
 168 explicitly forced ω_j (and not ω) to emphasize that the BSP is discrete from
 169 the outset, although nothing prevents us to write $U_d(\mathbf{r}, \omega)$ instead. However,
 170 this would give the false idea to the reader that in the $\sum_{j=1}^J$ -sum we are
 171 actually working with a BSP which is continuous in ω and then picking up
 172 specific contributions from $\omega = \omega_j$.

173 The generalized BSP can be split into TM and TE modes by simply
 174 appending lower indices “TM” and “TE”, respectively. At this point, the
 175 formulation can be carried out in one of the two ways. First, we use the
 176 TM and TE generalized BSPs and try to find the polychromatic, physical

177 fields. The second approach is to find the latter fields from the TM and TE
 178 monochromatic BSPs. As we will justify later, we prefer to follow the second
 179 path and reserve the superposition principle to be applied to the (physical)
 180 fields satisfying Maxwell's equation.

181 *2.1. Incident EM fields in the polychromatic GLMT*

182 From Maxwell's equations and the classes of orthogonal curvilinear coor-
 183 dinate systems constrained by Eq. (1), the following relations are established
 184 between EM fields $\mathbf{E}(\mathbf{r}, t)$ and $\mathbf{H}(\mathbf{r}, t)$, and the generalized BSP $U(\mathbf{r}, t)$, for
 185 both TM and TE modes, see Eqs. (2.23)–(2.35) in Ref. [28]:

$$E_{1, TM} = \partial_1^2 U_{TM} - \mu \epsilon \partial_t^2 U_{TM}, \quad (17)$$

$$E_{2, TM} = \frac{1}{e_2} \partial_1 \partial_2 U_{TM}, \quad (18)$$

$$E_{3, TM} = \frac{1}{e_3} \partial_1 \partial_3 U_{TM}, \quad (19)$$

$$H_{1, TM} = 0, \text{ (TM-definition),} \quad (20)$$

$$H_{2, TM} = \frac{\epsilon}{e_3} \partial_3 \partial_t U_{TM}, \quad (21)$$

$$H_{3, TM} = -\frac{\epsilon}{e_2} \partial_2 \partial_t U_{TM}, \quad (22)$$

191 for TM modes, and

$$E_{1, TE} = 0, \text{ (TE-definition),} \quad (23)$$

$$E_{2, TE} = -\frac{\mu}{e_3} \partial_3 \partial_t U_{TE}, \quad (24)$$

$$E_{3, TE} = \frac{\mu}{e_2} \partial_2 \partial_t U_{TE}, \quad (25)$$

$$H_{1, TE} = \partial_1^2 U_{TE} - \mu \epsilon \partial_t^2 U_{TE} \quad (26)$$

$$H_{2, TE} = \frac{1}{e_2} \partial_1 \partial_2 U_{TE}, \quad (27)$$

$$H_{3, TE} = \frac{1}{e_3} \partial_1 \partial_3 U_{TE} \quad (28)$$

197 for TE modes. In Eqs. (17)–(28), arguments have been omitted for simplicity,
 198 and the total electromagnetic field components are found from

$$E_k = E_{k, TM} + E_{k, TE}, \quad k = 1, 2, 3 \quad (29a)$$

$$H_k = H_{k,TM} + H_{k,TE}, \quad k = 1, 2, 3. \quad (29b)$$

As we mentioned in the previous section, using the superposition principle to define a generalized BSP is not very useful when $\mu \equiv \mu(\omega)$ and $\epsilon \equiv \epsilon(\omega)$, that is, when the constitutive parameters are frequency-dependent quantities in general. This can be readily seen from Eqs. (17), (21), (22), (24), (25) and (26), where multiplicative factors proportional to μ and ϵ are present which cannot be immediately absorbed inside the integral in Eq. (16) and become part of the integrand. Since it is \mathbf{E} and \mathbf{H} the physical fields of interest, we then first consider the monochromatic solutions to the electromagnetic field components and then use the superposition principle to establish their polychromatic, partial wave expressions.

Therefore, for the incident monochromatic fields, from Eq. (11) one readily writes,

$$U_{TX}^i(\mathbf{r}, t) = U_{TX}^i(\mathbf{r}, \omega) e^{i\omega t}, \quad (30)$$

where TX stands for either TM or TE .

Following the conventions in the monochromatic GLMT, let us further define:

$$c_{nm,TX} \equiv c_{nm,TX}(\omega) = c_n^{pw}(\omega) G_{n,TX}^m(\omega), \quad (31)$$

where

$$c_n^{pw}(\omega) = \frac{1}{ik(\omega)} (-i)^n \frac{2n+1}{n(n+1)}. \quad (32)$$

In Eq. (31), both field strengths E_0 and H_0 available from the monochromatic GLMT have been absorbed into $G_{n,TX}^m(\omega)$. Equation (32) can be seen as the natural generalization of c_n^{pw} in the monochromatic GLMT, see Eqs. (3.1)–(3.3) in Ref. [28], in order to deal with frequency-dependent fields.

Because Ricatti-Bessel functions $\xi_n(z)$ cannot be used to describe incident fields since it is not finite at the origin, from Eqs. (11), (31) and (32), the TM and TE instantaneous (complex) incident BSPs for monochromatic fields are then written as:

$$U_{TX}^i(\mathbf{r}, t) = \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} c_n^{pw}(\omega) G_{n,TX}^m(\omega) \frac{\psi_n[k(\omega)r]}{k(\omega)} P_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t}, \quad (33)$$

which obviously coincides with Eqs. (3.4) and (3.5) in Ref. [28], except for the fact that the dependence on ω is explicitly put into evidence. Therefore,

226 from the superposition principle, and making use of Eqs. (3.39)–(3.50) of
 227 Ref. [28], which corresponds to substituting Eq. (33) into (17)–(28) using
 228 spherical coordinates, one then obtains the general expressions for the partial
 229 wave expansion of polychromatic beams for TM modes:

$$\begin{aligned}
 E_{r,TM}^i(\mathbf{r}, t) &= \int_{-\infty}^{\infty} E_{r,TM}^i(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
 &= \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} k(\omega) c_n^{pw}(\omega) G_{n,TM}^m(\omega) \\
 &\quad \times [\psi_n''[k(\omega)r] + \psi_n[k(\omega)r]] P_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega,
 \end{aligned} \tag{34}$$

230

$$\begin{aligned}
 E_{\theta,TM}^i(\mathbf{r}, t) &= \int_{-\infty}^{\infty} E_{\theta,TM}^i(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
 &= \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} c_n^{pw}(\omega) G_{n,TM}^m(\omega) \frac{\psi_n'[k(\omega)r]}{r} \tau_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega,
 \end{aligned} \tag{35}$$

231

$$\begin{aligned}
 E_{\varphi,TM}^i(\mathbf{r}, t) &= \int_{-\infty}^{\infty} E_{\varphi,TM}^i(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
 &= i \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} m c_n^{pw}(\omega) G_{n,TM}^m(\omega) \frac{\psi_n'[k(\omega)r]}{r} \pi_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega,
 \end{aligned} \tag{36}$$

232

$$H_{r,TM}^i(\mathbf{r}, t) = \int_{-\infty}^{\infty} H_{r,TM}^i(\mathbf{r}, \omega) e^{i\omega t} d\omega = 0, \tag{37}$$

233

$$\begin{aligned}
 H_{\theta,TM}^i(\mathbf{r}, t) &= \int_{-\infty}^{\infty} H_{\theta,TM}^i(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
 &= - \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} m c_n^{pw}(\omega) \frac{G_{n,TM}^m(\omega)}{\eta(\omega)} \frac{\psi_n[k(\omega)r]}{r} \pi_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega,
 \end{aligned} \tag{38}$$

$$\begin{aligned}
H_{\varphi, TM}^i(\mathbf{r}, t) &= \int_{-\infty}^{\infty} H_{\varphi, TM}^i(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
&= -i \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} c_n^{pw}(\omega) \frac{G_{n, TM}^m(\omega)}{\eta(\omega)} \frac{\psi_n[k(\omega)r]}{r} \tau_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega.
\end{aligned} \tag{39}$$

234 Notice that Eqs. (34)-(36) are expressed in terms of $G_{n, TM}^m = E_0(\omega)g_{n, TM}^m$.
235 Therefore, in Eqs. (38) and (39), we have to introduce the intrinsic impedance
236 $\eta(\omega) = [\mu(\omega)/\epsilon(\omega)]^{1/2}$ so that, using Eq. (1.107) of Ref. [28], we have
237 $G_{n, TM}^m/\eta = H_0g_{n, TM}^m$. Note that, for similar reason, the impedance has to be
238 introduced in Eqs. (41) and (42) below.

239 Similarly, for TE modes,

$$E_{r, TE}^i(\mathbf{r}, t) = \int_{-\infty}^{\infty} E_{r, TE}^i(\mathbf{r}, \omega) e^{i\omega t} d\omega = 0, \tag{40}$$

240

$$\begin{aligned}
E_{\theta, TE}^i(\mathbf{r}, t) &= \int_{-\infty}^{\infty} E_{\theta, TE}^i(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
&= \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} m c_n^{pw}(\omega) \eta(\omega) G_{n, TE}^m(\omega) \frac{\psi_n[k(\omega)r]}{r} \pi_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega,
\end{aligned} \tag{41}$$

241

$$\begin{aligned}
E_{\varphi, TE}^i(\mathbf{r}, t) &= \int_{-\infty}^{\infty} E_{\varphi, TE}^i(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
&= i \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} c_n^{pw}(\omega) \eta(\omega) G_{n, TE}^m(\omega) \frac{\psi_n[k(\omega)r]}{r} \tau_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega,
\end{aligned} \tag{42}$$

242

$$\begin{aligned}
H_{r,TE}^i(\mathbf{r}, t) &= \int_{-\infty}^{\infty} H_{r,TE}^i(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
&= \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} k(\omega) c_n^{pw}(\omega) G_{n,TE}^m(\omega) \\
&\quad \times [\psi_n''[k(\omega)r] + \psi_n[k(\omega)r]] P_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega,
\end{aligned} \tag{43}$$

243

$$\begin{aligned}
H_{\theta,TE}^i(\mathbf{r}, t) &= \int_{-\infty}^{\infty} H_{\theta,TE}^i(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
&= \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} c_n^{pw}(\omega) G_{n,TE}^m(\omega) \frac{\psi_n'[k(\omega)r]}{r} \tau_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega,
\end{aligned} \tag{44}$$

244

$$\begin{aligned}
H_{\varphi,TE}^i(\mathbf{r}, t) &= \int_{-\infty}^{\infty} H_{\varphi,TE}^i(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
&= i \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} m c_n^{pw}(\omega) G_{n,TE}^m(\omega) \frac{\psi_n'[k(\omega)r]}{r} \pi_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega,
\end{aligned} \tag{45}$$

245 In Eqs. (34)–(45), $\pi_n^m(\cos \theta) = P_n^m(\cos \theta) / \sin \theta$ and $\tau_n^m(\cos \theta) = d_\theta P_n^m(\cos \theta)$
246 are generalized Legendre functions [28]. These equations are the polychromatic
247 generalization of the monochromatic Eqs. (3.39)–(3.50) in Ref. [28].

248 It is seen from Eqs. (34)–(45) that $G_{n,TE}^m(\omega)$ are frequency-dependent
249 functions. For discrete-frequency, polychromatic light, they are distributions
250 proportional to the BSCs of the original, monochromatic GLMT, as will soon
251 be shown. The proportionality depends on the field strengths $E_0(\omega_j)$ and
252 $H_0(\omega_j)$, respectively for electric and magnetic fields, of the j -th monochromatic
253 arbitrary-shaped beam.

254 For continuous frequency-dependent fields, $G_{n,TE}^m(\omega)$ are functions of a
255 continuous variable ω and represent FSSs. They become, therefore, a gener-
256 alization of the BSCs of the original GLMT, now extended to polychromatic
257 beams.

258 *2.2. Quadratures for the FSSs*

259 In this section we shall find general expressions to evaluate the FSSs
 260 $G_{n, TX}^m(\omega)$ by assuming arbitrary polychromatic fields which can be either
 261 discretely or continuously dependent on ω . To do so, orthogonal relations for
 262 the φ -exponential, $\exp(im\varphi)$, and associated Legendre functions, $P_n^{|m|}(\cos\theta)$,
 263 are imposed, as usual. Like in the monochromatic GLMT, it suffices to
 264 consider the radial field components $E_r^i(\mathbf{r}, t)$ and $H_r^i(\mathbf{r}, t)$ in Eq. (29), now
 265 with the lower index k of Eq. (29) representing the radial spherical coordinate
 266 r .

267 The total instantaneous radial field components $E_r^i(\mathbf{r}, t)$ and $H_r^i(\mathbf{r}, t)$ are
 268 found from Eqs. (29), (34), (37), (40) and (43):

$$\begin{aligned}
 E_r^i(\mathbf{r}, t) &= \int_{-\infty}^{\infty} E_r^i(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
 &= \int_{-\infty}^{\infty} \left\{ \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} n(n+1) c_n^{pw}(\omega) G_{n, TM}^m(\omega) \frac{\psi_n[k(\omega)r]}{k(\omega)r^2} P_n^{|m|}(\cos\theta) e^{im\varphi} \right\} e^{i\omega t} d\omega
 \end{aligned}
 \tag{46}$$

269 and

$$\begin{aligned}
 H_r^i(\mathbf{r}, t) &= \int_{-\infty}^{\infty} H_r^i(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
 &= \int_{-\infty}^{\infty} \left\{ \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} n(n+1) c_n^{pw}(\omega) G_{n, TE}^m(\omega) \frac{\psi_n[k(\omega)r]}{k(\omega)r^2} P_n^{|m|}(\cos\theta) e^{im\varphi} \right\} e^{i\omega t} d\omega,
 \end{aligned}
 \tag{47}$$

270 where we have used Ricatti-Bessel's differential equation $z^2\psi_n''(z) + [z^2 - n(n+1)]\psi_n(z) = 0$ to simplify the terms under brackets in Eqs. (34) and (43).

272 After introducing the following orthogonality relations for exponentials
 273 and associated Legendre functions [37]:

$$\int_0^{2\pi} \exp[i(m - m')\varphi] d\varphi = 2\pi\delta_{m, m'}, \tag{48}$$

274 and

$$\int_0^\pi P_n^{|m|}(\cos \theta) P_{n'}^{|m|}(\cos \theta) \sin \theta d\theta = \frac{2}{2n+1} \frac{(n+|m|)!}{(n-|m|)!} \delta_{n,n'}, \quad (49)$$

275 where $\delta_{i,j}$ is the Kronecker delta, Eqs. (46) and (47) can be recast under the
276 form:

$$\begin{aligned} & \int_{-\infty}^{\infty} c_n^{pw}(\omega) G_{n, TM}^m(\omega) \frac{\psi_n[k(\omega)r]}{k(\omega)} e^{i\omega t} d\omega \\ &= \frac{r^2}{4\pi} \frac{2n+1}{n(n+1)} \frac{(n-|m|)!}{(n+|m|)!} \int_0^\pi \int_0^{2\pi} E_r^i(\mathbf{r}, t) P_n^{|m|}(\cos \theta) e^{-im\varphi} \sin \theta d\varphi d\theta. \end{aligned} \quad (50)$$

277 and

$$\begin{aligned} & \int_{-\infty}^{\infty} c_n^{pw}(\omega) G_{n, TE}^m(\omega) \frac{\psi_n[k(\omega)r]}{k(\omega)} e^{i\omega t} d\omega \\ &= \frac{r^2}{4\pi} \frac{2n+1}{n(n+1)} \frac{(n-|m|)!}{(n+|m|)!} \int_0^\pi \int_0^{2\pi} H_r^i(\mathbf{r}, t) P_n^{|m|}(\cos \theta) e^{-im\varphi} \sin \theta d\varphi d\theta. \end{aligned} \quad (51)$$

278 Now, multiplying both sides of Eqs. (50) and (51) by $\exp(-i\omega't)$, inte-
279 grating over time, interchanging the order of the integrals in the l.h.s. of the
280 resulting expression and using the fact that, for continuous variables ω ,

$$\int_{-\infty}^{\infty} \exp[i(\omega - \omega')t] dt = 2\pi \delta(\omega - \omega'), \quad (52)$$

281 the following expressions for the FSSs $G_{n, TM}^m(\omega)$ and $G_{n, TE}^m(\omega)$ are then ob-
282 tained:

$$\begin{aligned} G_{n, TM}^m(\omega) &= \frac{r^2}{8\pi^2} \frac{2n+1}{c_n^{pw}(\omega)} \frac{(n-|m|)!}{n(n+1)(n+|m|)!} \frac{k(\omega)}{\psi_n[k(\omega)r]} \\ &\quad \times \int_0^\pi \int_0^{2\pi} \left[\int_{-\infty}^{\infty} E_r^i(\mathbf{r}, t) e^{-i\omega't} dt \right] P_n^{|m|}(\cos \theta) e^{-im\varphi} \sin \theta d\varphi d\theta \end{aligned} \quad (53)$$

283 and

$$\begin{aligned}
G_{n,TE}^m(\omega) &= \frac{r^2}{8\pi^2 c_n^{pw}(\omega)} \frac{2n+1}{n(n+1)} \frac{(n-|m|)!}{(n+|m|)!} \frac{k(\omega)}{\psi_n[k(\omega)r]} \\
&\times \int_0^\pi \int_0^{2\pi} \left[\int_{-\infty}^\infty H_r^i(\mathbf{r}, t) e^{-i\omega t} dt \right] P_n^{|m|}(\cos \theta) e^{-im\varphi} \sin \theta d\varphi d\theta.
\end{aligned} \tag{54}$$

284 Since the time integrals in Eqs. (53) and (54) correspond, in accordance
285 with Eq. (16) extended to the electric and magnetic fields, to 2π times the
286 Fourier transforms $E_r^i(\mathbf{r}, \omega)$ and $H_r^i(\mathbf{r}, \omega)$ of $E_r^i(\mathbf{r}, t)$ and $H_r^i(\mathbf{r}, t)$, respectively,
287 one finally gets the double quadrature formulae for the FSSs:

$$\begin{aligned}
G_{n,TM}^m(\omega) &= \frac{1}{4\pi c_n^{pw}(\omega)} \frac{2n+1}{n(n+1)} \frac{(n-|m|)!}{(n+|m|)!} \frac{k(\omega)r^2}{\psi_n[k(\omega)r]} \\
&\times \int_0^\pi \int_0^{2\pi} E_r^i(\mathbf{r}, \omega) P_n^{|m|}(\cos \theta) e^{-im\varphi} \sin \theta d\varphi d\theta
\end{aligned} \tag{55}$$

288 and

$$\begin{aligned}
G_{n,TE}^m(\omega) &= \frac{1}{4\pi c_n^{pw}(\omega)} \frac{2n+1}{n(n+1)} \frac{(n-|m|)!}{(n+|m|)!} \frac{k(\omega)r^2}{\psi_n[k(\omega)r]} \\
&\times \int_0^\pi \int_0^{2\pi} H_r^i(\mathbf{r}, \omega) P_n^{|m|}(\cos \theta) e^{-im\varphi} \sin \theta d\varphi d\theta.
\end{aligned} \tag{56}$$

289 Equations (55) and (56) represent the generalized quadrature procedure
290 to evaluate the FSSs in the polychromatic GLMT. In comparison with the
291 monochromatic GLMT, frequency-dependent quantities become explicitly
292 evident and, when no analytical solution exists to the time integrals in
293 Eqs. (53) and (54), an additional integral over time is present which must
294 be solved numerically. For electromagnetic fields constructed from discrete
295 superpositions of J monochromatic beams of distinct frequencies ω_j (for ex-
296 ample, a colorful intensity pattern created from a RGB laser system), ex-
297 pressing the physical, instantaneous fields in the form seen in the first term
298 of the r.h.s. of Eq. (14) allows us to write

$$E_r^i(\mathbf{r}, t) = \sum_{j=1}^J E_r^i(\mathbf{r}, \omega_j) e^{i\omega_j t} \tag{57}$$

$$H_r^i(\mathbf{r}, t) = \sum_{j=1}^J H_r^i(\mathbf{r}, \omega_j) e^{i\omega_j t}, \quad (58)$$

300 and, multiplying both sides of Eqs. (57) and (58) by $\exp(-i\omega t)$, integrating
301 with respect to time and making use of Eq. (52), one obtains:

$$\begin{aligned} \int_{-\infty}^{\infty} \left\{ \begin{array}{c} E_r^i(\mathbf{r}, t) \\ H_r^i(\mathbf{r}, t) \end{array} \right\} e^{-i\omega t} dt &= 2\pi \sum_{j=1}^J \left\{ \begin{array}{c} E_r^i(\mathbf{r}, \omega_j) \\ H_r^i(\mathbf{r}, \omega_j) \end{array} \right\} \delta(\omega - \omega_j) \\ &= 2\pi \left\{ \begin{array}{c} E_r^i(\mathbf{r}, \omega) \\ H_r^i(\mathbf{r}, \omega) \end{array} \right\} \end{aligned} \quad (59)$$

302 In Eqs. (57)–(59), $E_r^i(\mathbf{r}, \omega_j)$ and $H_r^i(\mathbf{r}, \omega_j)$ correspond to the electric and
303 magnetic field phasors of the monochromatic incident field of frequency $\omega =$
304 ω_j .

305 Therefore, from Eqs. (53), (54), (59),

$$\begin{aligned} G_{n, TM}^m(\omega) &= \sum_{j=1}^J \frac{1}{4\pi \mathcal{C}_n^{pw}(\omega)} \frac{2n+1}{n(n+1)} \frac{(n-|m|)!}{(n+|m|)!} \frac{k(\omega)r^2}{\psi_n[k(\omega)r]} \\ &\quad \times \int_0^\pi \int_0^{2\pi} E_r^i(\mathbf{r}, \omega_j) P_n^{|m|}(\cos\theta) e^{-im\varphi} \sin\theta d\varphi d\theta \delta(\omega - \omega_j) \\ &= \sum_{j=1}^J G_{n, TM}^m(\omega_j) \delta(\omega_j - \omega) \\ &= \sum_{j=1}^J E_0(\omega_j) g_{n, TM}^{m, mono}(\omega_j) \delta(\omega_j - \omega) \end{aligned} \quad (60)$$

306 and

$$\begin{aligned}
G_{n,TE}^m(\omega) &= \sum_{j=1}^J \frac{1}{4\pi c_n^{pw}(\omega)} \frac{2n+1}{n(n+1)} \frac{(n-|m|)!}{(n+|m|)!} \frac{k(\omega)r^2}{\psi_n[k(\omega)r]} \\
&\quad \times \int_0^\pi \int_0^{2\pi} H_r^i(\mathbf{r}, \omega) P_n^{|m|}(\cos\theta) e^{-im\varphi} \sin\theta d\varphi d\theta \delta(\omega - \omega_j) \\
&= \sum_{j=1}^J G_{n,TE}^m(\omega_j) \delta(\omega_j - \omega) \\
&= \sum_{j=1}^J H_0(\omega_j) g_{n,TE}^{m,mono}(\omega_j) \delta(\omega_j - \omega),
\end{aligned} \tag{61}$$

307 where, for discrete frequencies ω_j , $g_{n,TE}^{m,mono}(\omega_j)$ are the usual BSCs in the
308 monochromatic GLMT, $E_0(\omega_j)$ is a normalization electric field strength for
309 TM modes and $H_0(\omega_j)$ is a normalization magnetic field strength for TE
310 modes.

311 Equations (60) and (61) express the fact that the FSSs for discrete su-
312 perpositions of distinct, arbitrary-shaped time-harmonic fields, each of which
313 having a particular and arbitrary frequency ω_j , is the weighted sum of the
314 BSCs associated with each time-harmonic field. The sum is actually weighted
315 since $E_r^i(\mathbf{r}, \omega_j)$ and $H_r^i(\mathbf{r}, \omega_j)$ can possess different relative amplitudes. In
316 other words, the FSSs $G_{n,TE}^m(\omega)$ for discrete sums of monochromatic elec-
317 tromagnetic fields of distinct frequencies ω_j can be expressed as a train of
318 impulses at the discrete frequencies ω_j , each impulse weighted by a complex
319 amplitude $X_0(\omega_j)g_{n,TE}^m(\omega_j)$, in which $X_0(\omega_j) = E_0(\omega_j)$ for TM modes and
320 $X_0(\omega_j) = H_0(\omega_j)$ for TE modes. Equations (60) and (61), once substituted
321 in Eqs. (34)–(45), recover Eqs. (57) and (58), as expected.

322 For $\omega_j = \omega_0, \forall j$, that is, when all distinct, arbitrary-shaped fields carry
323 the same operating frequency (e.g., when generated by a single laser source),
324 the ω_j arguments can be dropped and the FSSs then become simple weighted
325 sums of equal-frequency BSCs representing a total monochromatic arbitrary-
326 shaped beam, such a problem lying therefore in the realm of the monochro-
327 matic GLMT.

328 Before ending this section, it must be stressed that, as it stands, the
329 monochromatic GLMT using BSPs expands the electromagnetic fields into
330 partial waves using spherical wave functions in spherical coordinates. This

331 leads to BSPs of the form given by Eq. (11) for a fixed but arbitrary ω ,
 332 where $U(\mathbf{r}, \omega)$ depends on both ω (an arbitrary but otherwise fixed frequency)
 333 and the spatial coordinates. However, it is common practice to express the
 334 instantaneous physical fields in terms of spectra of the form $S(\mathbf{k}, \omega)$, that is,
 335 by also transforming the spatial coordinates and going to a reciprocal space
 336 in the wave vector \mathbf{k} . For instance, paraxial, scalar Gaussian pulses can be
 337 represented by a spectrum $S(k_\rho, \omega)$, where k_ρ is the transverse wave number
 338 which obeys the dispersion relation $k^2 = \omega^2/c^2 = k_\rho^2 + k_z^2$, with k_z being the
 339 longitudinal wave number. Then, the corresponding scalar field $\psi(\rho, z, t)$ is
 340 determined from the following relation [38]:

$$\psi(\rho, z, t) = \int_0^\infty dk_\rho \int_{-\infty}^{+\infty} d\omega k_\rho S(k_\rho, \omega) J_0(k_\rho \rho) \exp\left(-i\sqrt{k^2 - k_\rho^2} z\right) e^{i\omega t}, \quad (62)$$

341 where $J_0(\cdot)$ is the zero-order cylindrical Bessel function of the first kind.

342 Extending this process to vector fields then would lead to another repre-
 343 sentation of Eqs. (55) and (56) involving integrations over k_ρ (or, in view of
 344 the dispersion relation, over k_z) with arguments involving spectral functions
 345 $E_r^i(\mathbf{k}, \omega)$ and $H_r^i(\mathbf{k}, \omega)$.

346 2.3. Scattered and internal fields

347 The procedure to find the scattered and internal field components resem-
 348 bles that adopted for the incident fields.

349 First, we introduce the instantaneous scattered and internal monochro-
 350 matic TM and TE BSPs in accordance with Eq. (30):

$$U_{TX}^s(\mathbf{r}, t) = U_{TX}^s(\mathbf{r}, \omega) e^{i\omega t}, \quad (63)$$

351

$$U_{TX}^{sp}(\mathbf{r}, t) = U_{TX}^{sp}(\mathbf{r}, \omega) e^{i\omega t}. \quad (64)$$

352 In Eqs. (63) and (64), the upper indices “s” and “sp” stands for scat-
 353 tered and internal (sphere) fields, respectively. In addition, $U_{TX}^s(\mathbf{r}, \omega)$ and
 354 $U_{TX}^{sp}(\mathbf{r}, \omega)$ are the BSP phasors in the monochromatic GLMT, see Eqs. (3.35)–
 355 (3.38), (3.84)–(3.87) of Ref. [28]. They are here reproduced for the conven-
 356 ience of the reader:

$$U_{TM}^s(\mathbf{r}, \omega) = - \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} \frac{1}{k} c_n^{pw} a_n G_{n, TM}^m \xi_n(kr) P_n^{|m|}(\cos \theta) e^{im\varphi}, \quad (65)$$

357

$$U_{TE}^s(\mathbf{r}, \omega) = - \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} \frac{1}{k} c_n^{pw} b_n G_{n,TE}^m \xi_n(kr) P_n^{|m|}(\cos \theta) e^{im\varphi}, \quad (66)$$

358

$$U_{TM}^{sp}(\mathbf{r}, \omega) = \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} \frac{k}{k_{sp}^2} c_n^{pw} c_n G_{n,TM}^m \psi_n(k_{sp}r) P_n^{|m|}(\cos \theta) e^{im\varphi}, \quad (67)$$

359

$$U_{TE}^{sp}(\mathbf{r}, \omega) = \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} \frac{k}{k_{sp}^2} c_n^{pw} d_n G_{n,TE}^m \psi_n(k_{sp}r) P_n^{|m|}(\cos \theta) e^{im\varphi}. \quad (68)$$

360 where a_n , b_n , c_n and d_n are the classical scattering coefficients of the Mie
361 theory [4, 28, 39]:

$$a_n = \frac{\mu_{sp} \psi_n(\alpha) \psi_n'(\beta) - \mu M \psi_n'(\alpha) \psi_n(\beta)}{\mu_{sp} \xi_n(\alpha) \psi_n'(\beta) - \mu M \xi_n'(\alpha) \psi_n(\beta)}, \quad (69)$$

362

$$b_n = \frac{\mu M \psi_n(\alpha) \psi_n'(\beta) - \mu_{sp} \psi_n'(\alpha) \psi_n(\beta)}{\mu M \xi_n(\alpha) \psi_n'(\beta) - \mu_{sp} \xi_n'(\alpha) \psi_n(\beta)}, \quad (70)$$

363

$$c_n = \frac{\mu_{sp} M [\xi_n(\alpha) \psi_n'(\alpha) - \xi_n'(\alpha) \psi_n(\alpha)]}{\mu_{sp} \xi_n(\alpha) \psi_n'(\beta) - \mu M \xi_n'(\alpha) \psi_n(\beta)}, \quad (71)$$

364

$$d_n = \frac{\mu M^2 [\xi_n(\alpha) \psi_n'(\alpha) - \xi_n'(\alpha) \psi_n(\alpha)]}{\mu M \xi_n(\alpha) \psi_n'(\beta) - \mu_{sp} \xi_n'(\alpha) \psi_n(\beta)}, \quad (72)$$

365 where

$$\alpha = ka = \frac{2\pi a}{\lambda} \quad (73a)$$

366

$$\beta = k_{sp}a = Mka = M\alpha. \quad (73b)$$

367 Notice that Eqs. (69)–(73) have been written for a specific but otherwise
368 arbitrary single operating frequency ω (arguments omitted). For a fixed
369 scatterer, the Mie coefficients are frequency-dependent functions, since μ_{sp} ,
370 M , α and β might also become functions of ω .

371 Substitution of Eqs. (63)–(68) into (17)–(28) and making use of the su-
372 perposition principle for the electromagnetic fields then leads to the scattered
373 and internal field components for TM and TE modes.

374

For the scattered fields, one obtains

$$\begin{aligned}
E_{r,TM}^s(\mathbf{r}, t) &= \int_{-\infty}^{\infty} E_{r,TM}^s(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
&= - \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} k(\omega) c_n^{pw}(\omega) a_n(\omega) G_{n,TM}^m(\omega) \\
&\quad \times [\xi_n''[k(\omega)r] + \xi_n[k(\omega)r]] P_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega,
\end{aligned} \tag{74}$$

375

$$\begin{aligned}
E_{\theta,TM}^s(\mathbf{r}, t) &= \int_{-\infty}^{\infty} E_{\theta,TM}^s(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
&= - \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} c_n^{pw}(\omega) a_n(\omega) G_{n,TM}^m(\omega) \frac{\xi_n'[k(\omega)r]}{r} \tau_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega,
\end{aligned} \tag{75}$$

376

$$\begin{aligned}
E_{\varphi,TM}^s(\mathbf{r}, t) &= \int_{-\infty}^{\infty} E_{\varphi,TM}^s(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
&= -i \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} m c_n^{pw}(\omega) a_n(\omega) G_{n,TM}^m(\omega) \frac{\xi_n[k(\omega)r]}{r} \pi_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega,
\end{aligned} \tag{76}$$

377

$$H_{r,TM}^s(\mathbf{r}, t) = \int_{-\infty}^{\infty} H_{r,TM}^s(\mathbf{r}, \omega) e^{i\omega t} d\omega = 0, \tag{77}$$

378

$$\begin{aligned}
H_{\theta,TM}^s(\mathbf{r}, t) &= \int_{-\infty}^{\infty} H_{\theta,TM}^s(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
&= \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} m c_n^{pw}(\omega) a_n(\omega) \frac{G_{n,TM}^m(\omega)}{\eta(\omega)} \frac{\xi_n[k(\omega)r]}{r} \pi_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega,
\end{aligned} \tag{78}$$

379

$$\begin{aligned}
H_{\varphi, TM}^s(\mathbf{r}, t) &= \int_{-\infty}^{\infty} H_{\varphi, TM}^s(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
&= i \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} c_n^{pw}(\omega) a_n(\omega) \frac{G_{n, TM}^m(\omega)}{\eta(\omega)} \frac{\xi_n[k(\omega)r]}{r} \tau_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega,
\end{aligned} \tag{79}$$

380 for TM modes, and

$$E_{r, TE}^s(\mathbf{r}, t) = \int_{-\infty}^{\infty} E_{r, TE}^s(\mathbf{r}, \omega) e^{i\omega t} d\omega = 0, \tag{80}$$

381

$$\begin{aligned}
E_{\theta, TE}^s(\mathbf{r}, t) &= \int_{-\infty}^{\infty} E_{\theta, TE}^s(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
&= - \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} m c_n^{pw}(\omega) b_n(\omega) \eta(\omega) G_{n, TE}^m(\omega) \frac{\xi_n[k(\omega)r]}{r} \tau_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega,
\end{aligned} \tag{81}$$

382

$$\begin{aligned}
E_{\varphi, TE}^s(\mathbf{r}, t) &= \int_{-\infty}^{\infty} E_{\varphi, TE}^s(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
&= -i \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} c_n^{pw}(\omega) b_n(\omega) \eta(\omega) G_{n, TE}^m(\omega) \frac{\xi_n[k(\omega)r]}{r} \tau_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega,
\end{aligned} \tag{82}$$

383

$$\begin{aligned}
H_{r, TE}^s(\mathbf{r}, t) &= \int_{-\infty}^{\infty} H_{r, TE}^s(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
&= - \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} k(\omega) c_n^{pw}(\omega) b_n(\omega) G_{n, TE}^m(\omega) \\
&\quad \times [\xi_n''[k(\omega)r] + \xi_n[k(\omega)r]] P_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega,
\end{aligned} \tag{83}$$

384

$$\begin{aligned}
H_{\theta,TE}^s(\mathbf{r}, t) &= \int_{-\infty}^{\infty} H_{\theta,TE}^s(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
&= - \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} c_n^{pw}(\omega) b_n(\omega) G_{n,TE}^m(\omega) \frac{\xi_n'[k(\omega)r]}{r} \tau_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega,
\end{aligned} \tag{84}$$

385

$$\begin{aligned}
H_{\varphi,TE}^s(\mathbf{r}, t) &= \int_{-\infty}^{\infty} H_{\varphi,TE}^s(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
&= -i \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} m c_n^{pw}(\omega) b_n(\omega) G_{n,TE}^m(\omega) \frac{\xi_n'[k(\omega)r]}{r} \pi_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega,
\end{aligned} \tag{85}$$

386 for TE modes.

387 Similarly, for the internal fields,

$$\begin{aligned}
E_{r,TM}^{sp}(\mathbf{r}, t) &= \int_{-\infty}^{\infty} E_{r,TM}^{sp}(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
&= \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} k(\omega) c_n^{pw}(\omega) c_n(\omega) G_{n,TM}^m(\omega) \\
&\quad \times [\psi_n''[k_{sp}(\omega)r] + \psi_n[k_{sp}(\omega)r]] P_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega,
\end{aligned} \tag{86}$$

388

$$\begin{aligned}
E_{\theta,TM}^{sp}(\mathbf{r}, t) &= \int_{-\infty}^{\infty} E_{\theta,TM}^{sp}(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
&= \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} \frac{k(\omega)}{k_{sp}(\omega)} c_n^{pw}(\omega) c_n(\omega) G_{n,TM}^m(\omega) \frac{\psi_n'[k_{sp}(\omega)r]}{r} \tau_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega,
\end{aligned} \tag{87}$$

389

$$\begin{aligned}
E_{\varphi, TM}^{sp}(\mathbf{r}, t) &= \int_{-\infty}^{\infty} E_{\varphi, TM}^{sp}(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
&= i \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} m \frac{k(\omega)}{k_{sp}(\omega)} c_n^{pw}(\omega) c_n(\omega) G_{n, TM}^m(\omega) \\
&\quad \times \frac{\psi'_n[k_{sp}(\omega)r]}{r} \pi_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega,
\end{aligned} \tag{88}$$

390

$$H_{r, TM}^{sp}(\mathbf{r}, t) = \int_{-\infty}^{\infty} H_{r, TM}^{sp}(\mathbf{r}, \omega) e^{i\omega t} d\omega = 0, \tag{89}$$

391

$$\begin{aligned}
H_{\theta, TM}^{sp}(\mathbf{r}, t) &= \int_{-\infty}^{\infty} H_{\theta, TM}^{sp}(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
&= - \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} m \frac{\mu(\omega)}{\mu_{sp}(\omega)} c_n^{pw}(\omega) c_n(\omega) \frac{G_{n, TM}^m(\omega)}{\eta(\omega)} \\
&\quad \times \frac{\psi_n[k_{sp}(\omega)r]}{r} \pi_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega,
\end{aligned} \tag{90}$$

392

$$\begin{aligned}
H_{\varphi, TM}^{sp}(\mathbf{r}, t) &= \int_{-\infty}^{\infty} H_{\varphi, TM}^{sp}(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
&= -i \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} \frac{\mu(\omega)}{\mu_{sp}(\omega)} c_n^{pw}(\omega) c_n(\omega) \frac{G_{n, TM}^m(\omega)}{\eta(\omega)} \\
&\quad \times \frac{\psi_n[k_{sp}(\omega)r]}{r} \tau_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega,
\end{aligned} \tag{91}$$

393 for TM modes, and

$$E_{r, TE}^{sp}(\mathbf{r}, t) = \int_{-\infty}^{\infty} E_{r, TE}^{sp}(\mathbf{r}, \omega) e^{i\omega t} d\omega = 0, \tag{92}$$

394

$$\begin{aligned}
E_{\theta,TE}^{sp}(\mathbf{r}, t) &= \int_{-\infty}^{\infty} E_{\theta,TE}^{sp}(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
&= \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} m \frac{\mu_{sp}(\omega)}{\mu(\omega)} \frac{k^2(\omega)}{k_{sp}^2(\omega)} c_n^{pw}(\omega) d_n(\omega) \eta(\omega) G_{n,TE}^m(\omega) \\
&\quad \times \frac{\psi_n[k_{sp}(\omega)r]}{r} \tau_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega,
\end{aligned} \tag{93}$$

395

$$\begin{aligned}
E_{\varphi,TE}^{sp}(\mathbf{r}, t) &= \int_{-\infty}^{\infty} E_{\varphi,TE}^{sp}(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
&= i \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} \frac{\mu_{sp}(\omega)}{\mu(\omega)} \frac{k^2(\omega)}{k_{sp}^2(\omega)} c_n^{pw}(\omega) d_n(\omega) \eta(\omega) G_{n,TE}^m(\omega) \\
&\quad \times \frac{\psi_n[k_{sp}(\omega)r]}{r} \tau_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega,
\end{aligned} \tag{94}$$

396

$$\begin{aligned}
H_{r,TE}^{sp}(\mathbf{r}, t) &= \int_{-\infty}^{\infty} H_{r,TE}^{sp}(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
&= \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} k(\omega) c_n^{pw}(\omega) d_n(\omega) G_{n,TE}^m(\omega) \\
&\quad \times [\psi_n''[k_{sp}(\omega)r] + \psi_n[k_{sp}(\omega)r]] P_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega,
\end{aligned} \tag{95}$$

397

$$\begin{aligned}
H_{\theta,TE}^{sp}(\mathbf{r}, t) &= \int_{-\infty}^{\infty} H_{\theta,TE}^{sp}(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
&= \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} \frac{k(\omega)}{k_{sp}(\omega)} c_n^{pw}(\omega) d_n(\omega) G_{n,TE}^m(\omega) \\
&\quad \times \frac{\psi_n'[k_{sp}(\omega)r]}{r} \tau_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega,
\end{aligned} \tag{96}$$

$$\begin{aligned}
H_{\varphi,TE}^{sp}(\mathbf{r}, t) &= \int_{-\infty}^{\infty} H_{\varphi,TE}^{sp}(\mathbf{r}, \omega) e^{i\omega t} d\omega \\
&= i \int_{-\infty}^{\infty} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} m \frac{k(\omega)}{k_{sp}(\omega)} c_n^{pw}(\omega) d_n(\omega) G_{n,TE}^m(\omega) \\
&\quad \times \frac{\psi'_n[k_{sp}(\omega)r]}{r} \pi_n^{|m|}(\cos \theta) e^{im\varphi} e^{i\omega t} d\omega,
\end{aligned} \tag{97}$$

399 for TE modes.

400 3. Physical quantities in the polychromatic GLMT

401 In the monochromatic GLMT, several physical quantities of practical im-
402 portance can be expressed in terms of the usual BSCs. Among them we find,
403 for instance, scattered intensities, phase angles, absorption, scattering and
404 extinction cross sections, radiation pressure forces and torques.

405 It is instructive to investigate the changes occurred in such quantities and
406 re-express them when one goes from single time-harmonic to arbitrary time-
407 dependent fields. It must be stressed that, since physical quantities in the
408 monochromatic GLMT are calculated from time-averaged fields, a general
409 procedure for extending them to polychromatic fields can be devised, as we
410 now show.

411 3.1. Scattered Intensities

412 The scattered intensities are found from time-averaging, for the far-fields,
413 the instantaneous Poynting vector $\mathbf{S}(\mathbf{r}, t) = \mathbf{E}^s(\mathbf{r}, t) \times \mathbf{H}^s(\mathbf{r}, t)$. Because
414 instantaneous fields are real quantities, one can use the relations $\mathbf{E}^s(\mathbf{r}, t) =$
415 $[\mathbf{E}^s(\mathbf{r}, t) + \mathbf{E}^{s*}(\mathbf{r}, t)]/2$ and $\mathbf{H}^s(\mathbf{r}, t) = [\mathbf{H}^s(\mathbf{r}, t) + \mathbf{H}^{s*}(\mathbf{r}, t)]/2$ so that, after
416 some rearrangement, it follows that:

$$\mathbf{S}(\mathbf{r}, t) = \frac{1}{2} \text{Re} [\mathbf{E}^s(\mathbf{r}, t) \times \mathbf{H}^s(\mathbf{r}, t)] + \frac{1}{2} \text{Re} [\mathbf{E}^s(\mathbf{r}, t) \times \mathbf{H}^{s*}(\mathbf{r}, t)]. \tag{98}$$

417 From the first lines of Eqs. (34)–(45), (74)–(85) and (86)–(97), one writes:

$$\mathbf{E}^k(\mathbf{r}, t) = \int_{-\infty}^{\infty} \mathbf{E}^k(\mathbf{r}, \omega) e^{i\omega t} d\omega, \tag{99}$$

$$\mathbf{H}^k(\mathbf{r}, t) = \int_{-\infty}^{\infty} \mathbf{H}^k(\mathbf{r}, \omega) e^{i\omega t} d\omega, \quad (100)$$

419 where $k = i, s, sp$ for incident, scattered and sphere fields, respectively.

420 In view of Eqs. (99) and (100), the cross product $\mathbf{E}^s(\mathbf{r}, t) \times \mathbf{H}^s(\mathbf{r}, t)$ in
421 Eq. (98) can be expressed as:

$$\mathbf{E}^s(\mathbf{r}, t) \times \mathbf{H}^s(\mathbf{r}, t) = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \mathbf{E}^s(\mathbf{r}, \omega) \times \mathbf{H}^s(\mathbf{r}, \omega') e^{i(\omega+\omega')t} d\omega d\omega' \quad (101)$$

422 Once averaged over time, the term in the r.h.s. of Eq. (101) will be
423 proportional to an integral of the form

$$\lim_{T \rightarrow \infty} \frac{1}{T} \int_{-T/2}^{T/2} e^{i(\omega+\omega')t} dt = \begin{cases} 1, & \omega = -\omega' \\ 0, & \text{otherwise} \end{cases}. \quad (102)$$

424 For scattered fields with non-negative frequencies, the cross product $\mathbf{E}^s(\mathbf{r}, \omega) \times$
425 $\mathbf{H}^s(\mathbf{r}, \omega)$ will contain only positive frequency components, that is, ω and ω'
426 will both be positive. Therefore, the use of Eq. (102) together with the fact
427 that $\omega > 0$ and $\omega' > 0$, imply that Eq. (101) will average to zero.

428 Now, for the second term in the r.h.s. of Eq. (98), it is seen that

$$\mathbf{E}^s(\mathbf{r}, t) \times \mathbf{H}^{s*}(\mathbf{r}, t) = \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \mathbf{E}^s(\mathbf{r}, \omega) \times \mathbf{H}^{s*}(\mathbf{r}, \omega') e^{i(\omega-\omega')t} d\omega d\omega'. \quad (103)$$

429 Similarly, since

$$\lim_{T \rightarrow \infty} \frac{1}{T} \int_{-T/2}^{T/2} e^{i(\omega-\omega')t} dt = \begin{cases} 1, & \omega = \omega' \\ 0, & \text{otherwise} \end{cases}, \quad (104)$$

430 for physical scattered fields carrying no non-negative frequencies, once aver-
431 aged over time, the term in the r.h.s. of Eq. (103) will simplify to a single

432 integral (either in ω or ω') and, using Eq. (98), we will eventually be left
 433 with:

$$\begin{aligned}
 \langle \mathbf{S}(\mathbf{r}, t) \rangle &= \lim_{T \rightarrow \infty} \frac{1}{T} \int_{-T/2}^{+T/2} \mathbf{S}(\mathbf{r}, t) dt \\
 &= \left\langle \frac{1}{2} \operatorname{Re} [\mathbf{E}^s(\mathbf{r}, t) \times \mathbf{H}^{s*}(\mathbf{r}, t)] \right\rangle \\
 &= \frac{1}{2} \operatorname{Re} \left\{ \int_{-\infty}^{\infty} \mathbf{E}^s(\mathbf{r}, \omega) \times \mathbf{H}^{s*}(\mathbf{r}, \omega) d\omega \right\},
 \end{aligned} \tag{105}$$

434 or, more compactly,

$$\langle \mathbf{S}(\mathbf{r}, t) \rangle = \operatorname{Re} \left[\int_{-\infty}^{+\infty} \mathbf{S}(\mathbf{r}, \omega) d\omega \right]. \tag{106}$$

435 with

$$\mathbf{S}(\mathbf{r}, \omega) = \frac{1}{2} \mathbf{E}^s(\mathbf{r}, \omega) \times \mathbf{H}^{s*}(\mathbf{r}, \omega). \tag{107}$$

436 From Eqs. (74)–(85), after imposing the far-field asymptotic expressions
 437 [28]

$$\xi_n(z) \rightarrow i^{n+1} e^{-ikr}, \tag{108a}$$

$$\xi_n''(z) + \xi_n(z) = 0, \tag{108b}$$

439 one then has the following non-normalized scattered intensity for polychro-
 440 matic fields:

$$\begin{aligned}
 S_c &\equiv S_c(\omega) = \operatorname{Re} [\mathbf{S}(\mathbf{r}, \omega)] \\
 &= \frac{1}{2} \operatorname{Re} [E_\theta^s(\mathbf{r}, \omega) H_\varphi^{s*}(\mathbf{r}, \omega) - E_\varphi^s(\mathbf{r}, \omega) H_\theta^{s*}(\mathbf{r}, \omega)] \\
 &= \frac{1}{2} \frac{1}{\eta(\omega) k^2(\omega) r^2} (|\mathcal{S}_{2,c}|^2 + |\mathcal{S}_{1,c}|^2) \\
 &= \frac{1}{2\eta(\omega) k^2(\omega) r^2} |\mathcal{S}_{2,c}|^2 + \frac{1}{2\eta(\omega) k^2(\omega) r^2} |\mathcal{S}_{1,c}|^2 \\
 &= I_{\theta,c} + I_{\varphi,c},
 \end{aligned} \tag{109}$$

441 where

$$I_{\theta,c} \equiv I_{\theta,c}(\omega) = \frac{1}{2\eta(\omega)k^2(\omega)r^2} |\mathcal{S}_{2,c}|^2 \quad (110)$$

442 and

$$I_{\varphi,c} \equiv I_{\varphi,c}(\omega) = \frac{1}{2\eta(\omega)k^2(\omega)r^2} |\mathcal{S}_{1,c}|^2 \quad (111)$$

443 In Eqs. (109)–(111), the $1/2\eta(\omega)$ comes from the fact that the normal-
444 ization condition of Eq. (3.106) of Ref. [28] was not considered. Also, the
445 generalized amplitude functions $\mathcal{S}_{1,c}(\omega)$ and $\mathcal{S}_{2,c}(\omega)$ are given by:

$$\begin{aligned} \mathcal{S}_{1,c} \equiv \mathcal{S}_{1,c}(\omega) = & \sum_{n=1}^{\infty} \sum_{m=-n}^n \frac{2n+1}{n(n+1)} \left[ma_n(\omega) G_{n,TM}^m(\omega) \pi_n^{|m|}(\cos\theta) \right. \\ & \left. + ib_n(\omega) \eta(\omega) G_{n,TE}^m(\omega) \tau_n^{|m|}(\cos\theta) \right] e^{im\varphi}, \end{aligned} \quad (112)$$

446

$$\begin{aligned} \mathcal{S}_{2,c} \equiv \mathcal{S}_{2,c}(\omega) = & \sum_{n=1}^{\infty} \sum_{m=-n}^n \frac{2n+1}{n(n+1)} \left[a_n(\omega) G_{n,TM}^m(\omega) \tau_n^{|m|}(\cos\theta) \right. \\ & \left. + imb_n(\omega) \eta(\omega) G_{n,TE}^m(\omega) \pi_n^{|m|}(\cos\theta) \right] e^{im\varphi}. \end{aligned} \quad (113)$$

447 In Eqs. (112) and (113), the FSSs $G_{n,TM}^m(\omega)$ and $G_{n,TE}^m(\omega)$ can represent
448 discrete or continuous frequency-dependent incident fields, or both. Notice
449 that the second line of Eq. (109) corresponds to Eq. (3.105) of Ref. [28] for
450 fixed but otherwise arbitrary positive frequencies, and that Eqs. (112) and
451 (113) re-express their monochromatic versions as given by Eqs. (3.110) and
452 (3.111) of Ref. [28], now in terms of the FSSs $G_{n,TM}^m(\omega)$ and $G_{n,TE}^m(\omega)$.

453 In conclusion, in view of Eqs. (106), (109)–(113), the total scattered in-
454 tensity, S , can be expressed as:

$$S = \int_{-\infty}^{\infty} S_c d\omega = \int_{-\infty}^{\infty} [I_{\theta,c} + I_{\varphi,c}] d\omega = I_{\theta} + I_{\varphi}, \quad (114)$$

455 where

$$\begin{bmatrix} I_{\theta} \\ I_{\varphi} \end{bmatrix} = \frac{1}{r^2} \int_{-\infty}^{\infty} \begin{bmatrix} \frac{|\mathcal{S}_{2,c}|^2}{2\eta(\omega)k^2(\omega)} \\ \frac{|\mathcal{S}_{1,c}|^2}{2\eta(\omega)k^2(\omega)} \end{bmatrix} d\omega \quad (115)$$

456 This extends the scattered intensities in the monochromatic GLMT to
457 the more general case of arbitrary frequency-dependent fields in the poly-
458 chromatic GLMT.

459 *3.2. Phase angles*

460 Phase angles can be used to describe the state of polarization of a given
 461 scattered electromagnetic field in the far field. Since the angular field com-
 462 ponents are the only non-negligible components in the far field, the angle
 463 associated with the ratio E_θ^s/E_φ^s can be used as a measure of such polariza-
 464 tion.

465 To see it clearly, consider Eqs. (74)–(85) in the far field, that is, with
 466 the asymptotic expressions given by Eq. (108). Following usual notation
 467 presented, e.g., in Sec. 3.10 of Ref. [28], omitting arguments and making use
 468 of Eqs. (32), (112) and (113), one finds:

$$E_\theta^s = i \int_{-\infty}^{\infty} \frac{e^{-ikr}}{kr} \mathcal{S}_{2,c} d\omega = \mathcal{E}_\theta, \quad (116a)$$

$$E_\varphi^s = - \int_{-\infty}^{\infty} \frac{e^{-ikr}}{kr} \mathcal{S}_{1,c} d\omega = \mathcal{E}_\varphi, \quad (116b)$$

470 In Eq. (116),

$$\mathcal{E}_\theta = i \int_{-\infty}^{\infty} \frac{e^{-ikr}}{kr} \mathcal{S}_{2,c} d\omega = i\mathcal{S}_2, \quad (117a)$$

$$\mathcal{E}_\varphi = - \int_{-\infty}^{\infty} \frac{e^{-ikr}}{kr} \mathcal{S}_{1,c} d\omega = -\mathcal{S}_1, \quad (117b)$$

472 with \mathcal{S}_1 and \mathcal{S}_2 being the non-normalized, total generalized amplitude func-
 473 tions.

474 Let δ denote the phase angle between the total field components E_θ^s and
 475 E_φ^s , which may be used to characterize the state of polarization of the total
 476 polychromatic beam. Then, from Eq. (117) and defining

$$\mathcal{E}_\theta = i\mathcal{S}_2 = A_\theta e^{i\varphi_2} \quad (118a)$$

$$\mathcal{E}_\varphi = -\mathcal{S}_1 = A_\varphi e^{i\varphi_1}, \quad (118b)$$

478 where A_θ and A_φ are real amplitude functions and φ_2 and φ_1 are the respec-
 479 tive phase angles of \mathcal{E}_θ and \mathcal{E}_φ , it is then seen that

$$\begin{aligned} \tan \delta &= \tan(\varphi_2 - \varphi_1) \\ &= \frac{\operatorname{Re}(\mathcal{S}_1) \operatorname{Re}(\mathcal{S}_2) + \operatorname{Im}(\mathcal{S}_1) \operatorname{Im}(\mathcal{S}_2)}{\operatorname{Im}(\mathcal{S}_1) \operatorname{Re}(\mathcal{S}_2) - \operatorname{Re}(\mathcal{S}_1) \operatorname{Im}(\mathcal{S}_2)}. \end{aligned} \quad (119)$$

480 Equation (119) is identical in form with the one derived in the monochro-
 481 matic GLMT, see Eq. (3.114) in Ref. [28]. However, now \mathcal{S}_1 and \mathcal{S}_2 are
 482 defined in accordance with Eq. (117), that is, they carry the spherical wave
 483 factor $\exp(-ikr)/kr$ inside the integral over ω .

484 When the incident field is of the form given by Eqs. (57) and (58), substi-
 485 tution of the corresponding FSSs of Eqs. (60) and (61) into (117), with the
 486 aid of (112) and (113), leads to the simplified forms

$$\begin{aligned}
 \mathcal{S}_1 &= \int_{-\infty}^{\infty} \frac{e^{-ik(\omega)r}}{k(\omega)r} \mathcal{S}_{1,c}(\omega) d\omega \\
 &= \int_{-\infty}^{\infty} \frac{e^{-ik(\omega)r}}{k(\omega)r} \sum_{j=1}^J E_0(\omega_j) \mathcal{S}_1^{mono}(\omega_j) \delta(\omega - \omega_j) d\omega \\
 &= \sum_{j=1}^J \frac{e^{-ik(\omega_j)r}}{k(\omega_j)r} E_0(\omega_j) \mathcal{S}_1^{mono}(\omega_j)
 \end{aligned} \tag{120}$$

487 and

$$\begin{aligned}
 \mathcal{S}_2 &= \int_{-\infty}^{\infty} \frac{e^{-ik(\omega)r}}{k(\omega)r} \mathcal{S}_{2,c}(\omega) d\omega \\
 &= \int_{-\infty}^{\infty} \frac{e^{-ik(\omega)r}}{k(\omega)r} \sum_{j=1}^J E_0(\omega_j) \mathcal{S}_2^{mono}(\omega_j) \delta(\omega - \omega_j) d\omega \\
 &= \sum_{j=1}^J \frac{e^{-ik(\omega_j)r}}{k(\omega_j)r} E_0(\omega_j) \mathcal{S}_2^{mono}(\omega_j),
 \end{aligned} \tag{121}$$

488 where $\mathcal{S}_1^{mono}(\omega_j) = \mathcal{S}_{1,c}(\omega = \omega_j)/E_0(\omega_j)$ and $\mathcal{S}_2^{mono}(\omega_j) = \mathcal{S}_{2,c}(\omega = \omega_j)/E_0(\omega_j)$
 489 are the normalized, generalized amplitude functions in the monochromatic
 490 GLMT, see also Eqs. (3.106), (3.110) and (3.111) of Ref. [28], for a particular
 491 operating frequency ω_j . Obviously, for $J = 1$, the identification between the
 492 polychromatic and the monochromatic GLMT is complete, and the phase
 493 angle δ as given by Eq. (119) will coincide with that in the latter framework,
 494 as it should.

495 3.3. Extinction, scattering and absorption cross-sections

496 Other physical quantities of interest in the polychromatic GLMT include
 497 the scattering, extinction and absorbing cross sections \mathcal{C}_{sca} , \mathcal{C}_{ext} and \mathcal{C}_{abs} ,

498 respectively, which are calculated by considering the total radial Poynting
 499 vector and integrating it over a spherical surface centered at the origin of the
 500 coordinate system (where the center of the scatterer is located).

501 The results are straightforward, but we shall present them for clarity.
 502 Let C_k^{mono} ($k = abs, sca$ or ext) denote the normalized cross sections familiar
 503 from the monochromatic GLMT, see Eqs.(3.106), (3.124), (3.137) and (3.142)
 504 in Ref. [28]. Proceeding similarly to Sec. 3.1, one eventually establishes that,
 505 for arbitrary-shaped, polychromatic fields in free space,

$$C_{abs} = C_{ext} - C_{sca} \quad (122)$$

506

$$\begin{aligned} C_{sca} &= \int_{-\infty}^{\infty} \frac{|E_0(\omega)|^2}{2\eta(\omega)} C_{sca}^{mono}(\omega) d\omega \\ &= \frac{1}{\pi} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} \frac{2n+1}{n(n+1)} \frac{(n+|m|)!}{(n-|m|)!} \\ &\quad \times \int_{-\infty}^{\infty} \frac{\lambda^2(\omega)}{2\eta(\omega)} \left(|a_n(\omega)|^2 |G_{n,TM}^m(\omega)|^2 + |b_n(\omega)|^2 \eta^2(\omega) |G_{n,TE}^m(\omega)|^2 \right) d\omega, \end{aligned} \quad (123)$$

507

$$\begin{aligned} C_{ext} &= \int_{-\infty}^{\infty} \frac{|E_0(\omega)|^2}{2\eta(\omega)} C_{ext}^{mono}(\omega) d\omega \\ &= \frac{1}{\pi} \operatorname{Re} \left[\sum_{n=1}^{\infty} \sum_{m=-n}^{+n} \frac{2n+1}{n(n+1)} \frac{(n+|m|)!}{(n-|m|)!} \right. \\ &\quad \times \left. \int_{-\infty}^{\infty} \frac{\lambda^2(\omega)}{2\eta(\omega)} \left(a_n(\omega) |G_{n,TM}^m(\omega)|^2 + b_n(\omega) \eta^2(\omega) |G_{n,TE}^m(\omega)|^2 \right) d\omega \right], \end{aligned} \quad (124)$$

508 with simplifications again being possible for discrete sums of monochromatic
 509 fields of different or equal frequencies, further simplified for single monochro-
 510 matic fields.

511 3.4. Radiation pressure cross-sections

512 As a final example of a physical quantity extended to polychromatic fields,
 513 we consider radiation pressure cross sections $C_{pr,i}$ in Cartesian coordinates

514 ($i = x, y, z$). In the monochromatic GLMT, such radiation pressure cross-
 515 sections can be found, e.g., in Eqs. (3.159), (3.181) and (3.185) of Ref. [28].

516 Assuming that the incident light propagates along $+z$, as usual, in the
 517 polychromatic GLMT the new $C_{pr,i}$ will read as:

$$\begin{aligned}
 C_{pr,z} = & \frac{1}{\pi} \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} \frac{1}{(n+1)^2} \frac{(n+1+|m|)!}{(n-|m|)!} \int_{-\infty}^{\infty} \frac{\lambda^2(\omega)}{2\eta(\omega)} \\
 & \times \left\{ \text{Re} \left[(a_n(\omega) + a_{n+1}^*(\omega) - 2a_n(\omega)a_{n+1}^*(\omega)) G_{n,TM}^m(\omega) G_{n+1,TM}^{m*}(\omega) \right. \right. \\
 & + (b_n(\omega) + b_{n+1}^*(\omega) - 2b_n(\omega)b_{n+1}^*(\omega)) \eta^2(\omega) G_{n,TE}^m(\omega) G_{n+1,TE}^{m*}(\omega) \left. \left. \right] \right. \\
 & + m \frac{2n+1}{n^2(n+1)^2} \frac{(n+|m|)!}{(n-|m|)!} \\
 & \left. \times \text{Re} \left[i(2a_n(\omega)b_n^*(\omega) - a_n(\omega) - b_n^*(\omega)) G_{n,TM}^m(\omega) \eta(\omega) G_{n,TE}^{m*}(\omega) \right] \right\} d\omega
 \end{aligned} \tag{125}$$

518 for the longitudinal component, and

$$\begin{aligned}
 \left\{ \begin{array}{l} C_{pr,x} \\ C_{pr,y} \end{array} \right\} = & \frac{1}{2\pi} \sum_{p=1}^{\infty} \sum_{n=p}^{\infty} \sum_{m=p-1 \neq 0}^{\infty} \frac{(n+p)!}{(n-p)!} \\
 & \times \left\{ \left(\frac{1}{m^2} \delta_{m,n+1} - \frac{1}{n^2} \delta_{n,m+1} \right) \right. \\
 & \times \int_{-\infty}^{\infty} \frac{\lambda^2(\omega)}{2\eta(\omega)} \left\{ \begin{array}{l} \text{Re} \\ \text{Im} \end{array} \right\} [S_{mn}^{p-1}(\omega) + S_{nm}^{-p}(\omega) - 2U_{mn}^{p-1}(\omega) - 2U_{nm}^{-p}(\omega)] d\omega \\
 & + \frac{2n+1}{n^2(n+1)^2} \delta_{n,m} \\
 & \left. \times \int_{-\infty}^{\infty} \frac{\lambda^2(\omega)}{2\eta(\omega)} \left\{ \begin{array}{l} \text{Re} \\ \text{Im} \end{array} \right\} [T_{mn}^{p-1}(\omega) - T_{nm}^{-p}(\omega) - 2V_{mn}^{p-1}(\omega) + 2V_{nm}^{-p}(\omega)] d\omega \right\}
 \end{aligned} \tag{126}$$

519 In Eqs. (125) and (126), the expressions for S_{nm}^p , T_{nm}^p , U_{nm}^p and V_{nm}^p are
 520 similar to those available from the monochromatic GLMT [see Eqs. (3.167),

521 (3.168), (3.177) and (3.178) in Ref. [28], a typo being corrected in the last
 522 term of Eq. (3.178)]:

$$S_{nm}^p(\omega) = [a_n(\omega) + a_m^*(\omega)] G_{n,TM}^p(\omega) G_{m,TM}^{p+1*}(\omega) + [b_n(\omega) + b_m^*(\omega)] \eta^2(\omega) G_{n,TE}^p(\omega) G_{m,TE}^{p+1*}(\omega), \quad (127)$$

523

$$T_{nm}^p(\omega) = -i [a_n(\omega) + b_m^*(\omega)] G_{n,TM}^p(\omega) \eta(\omega) G_{m,TE}^{p+1*}(\omega) + i [b_n(\omega) + a_m^*(\omega)] \eta(\omega) G_{n,TE}^p(\omega) G_{m,TM}^{p+1*}(\omega), \quad (128)$$

524

$$U_{nm}^p(\omega) = a_n(\omega) a_m^*(\omega) G_{n,TM}^p(\omega) G_{m,TM}^{p+1*}(\omega) + b_n(\omega) b_m^*(\omega) \eta^2(\omega) G_{n,TE}^p(\omega) G_{m,TE}^{p+1*}(\omega), \quad (129)$$

525

$$V_{nm}^p(\omega) = i b_n(\omega) a_m^*(\omega) \eta(\omega) G_{n,TE}^p(\omega) G_{m,TM}^{p+1*}(\omega) - i a_n(\omega) b_m^*(\omega) G_{n,TM}^p(\omega) \eta(\omega) G_{m,TE}^{p+1*}(\omega). \quad (130)$$

526 For incident fields generated from discrete superpositions of monochromatic
 527 fields of arbitrary frequencies ω_j , substitution of Eqs. (60) and (61)
 528 into (125) and (126) allows us to perform the integral over ω and get

$$C_{pr,z} = \frac{1}{\pi} \sum_{j=1}^J \sum_{n=1}^{\infty} \sum_{m=-n}^{+n} \frac{1}{(n+1)^2} \frac{(n+1+|m|)! \lambda^2(\omega_j)}{(n-|m|)! 2\eta(\omega_j)} \times \left\{ \begin{aligned} & \text{Re} \left[(a_n(\omega_j) + a_{n+1}^*(\omega_j) - 2a_n(\omega_j) a_{n+1}^*(\omega_j)) |E_0(\omega_j)|^2 g_{n,TM}^{m,mono}(\omega_j) g_{n+1,TM}^{m*,mono}(\omega_j) \right. \\ & + (b_n(\omega_j) + b_{n+1}^*(\omega_j) - 2b_n(\omega_j) b_{n+1}^*(\omega_j)) |E_0(\omega_j)|^2 g_{n,TE}^{m,mono}(\omega_j) g_{n+1,TE}^{m*,mono}(\omega_j) \\ & + m \frac{2n+1}{n^2(n+1)^2} \frac{(n+|m|)!}{(n-|m|)!} \\ & \left. \times \text{Re} \left[i(2a_n(\omega_j) b_n^*(\omega_j) - a_n(\omega_j) - b_n^*(\omega_j)) |E_0(\omega_j)|^2 g_{n,TM}^{m,mono}(\omega_j) g_{n,TE}^{m*,mono}(\omega_j) \right] \right\} \quad (131)$$

$$\begin{aligned}
\left\{ \begin{array}{l} C_{pr,x} \\ C_{pr,y} \end{array} \right\} &= \frac{1}{2\pi} \sum_{j=1}^J \sum_{p=1}^{\infty} \sum_{n=p}^{\infty} \sum_{m=p-1 \neq 0}^{\infty} \frac{(n+p)! \lambda^2(\omega_j)}{(n-p)! 2\eta(\omega_j)} \\
&\times \left\{ \left(\frac{1}{m^2} \delta_{m,n+1} - \frac{1}{n^2} \delta_{n,m+1} \right) \right. \\
&\left. \left\{ \begin{array}{l} \text{Re} \\ \text{Im} \end{array} \right\} [S_{mn}^{p-1}(\omega_j) + S_{nm}^{-p}(\omega_j) - 2U_{mn}^{p-1}(\omega_j) - 2U_{nm}^{-p}(\omega_j)] \right. \\
&+ \frac{2n+1}{n^2(n+1)^2} \delta_{n,m} \\
&\left. \times \left\{ \begin{array}{l} \text{Re} \\ \text{Im} \end{array} \right\} [T_{mn}^{p-1}(\omega_j) - T_{nm}^{-p}(\omega_j) - 2V_{mn}^{p-1}(\omega_j) + 2V_{nm}^{-p}(\omega_j)] \right\} \\
&\hspace{15em} (132)
\end{aligned}$$

530 with the FFSs in Eqs. (127)–(130) being now replaced by $X_0 g_{n,TX}^{m,mono}(\omega_j)$. If
531 $\omega_j = \omega_0, \forall j$, then the total field is a monochromatic field with operating fre-
532 quency ω_0 , therefore falling into the realm of the monochromatic GLMT. Of
533 course, for a single monochromatic incident field, the j -sum can be discarded,
534 and the radiation pressure cross-sections become the familiar cross-sections
535 in the monochromatic GLMT. Torque expressions for polychromatic fields
536 may be treated similarly following Refs. [40, 41].

537 4. Conclusions

538 In this work, a polychromatic version of the generalized Lorenz-Mie theory
539 has been proposed in order to describe the interaction between an arbitrary-
540 shaped, arbitrarily frequency-dependent optical field and a spherical, homo-
541 geneous particle, in terms of a partial wave expansion starting from Bromwich
542 scalar potentials.

543 The monochromatic Bromwich scalar potentials have been expanded into
544 spherical wave functions and their instantaneous, complex forms used to de-
545 rive expressions for the incident electromagnetic fields. It was shown that
546 the total incident field can be described in terms of field shape spectra, since
547 frequency now becomes, in general, a continuous variable. When the total in-
548 cident field becomes a discrete superposition of arbitrary-shaped, monochro-
549 matic light waves, the field shape spectra reduces to a train of impulses

550 whose complex amplitude associated with each impulse is the beam shape
551 coefficient of the original, monochromatic generalized Lorenz-Mie theory.

552 Expressions for several physical quantities of interest have been provided.
553 They include incident, scattered and internal fields, scattered intensities and
554 phase angles, extinction, scattering and absorption cross-sections, and ra-
555 diation pressure cross-sections which are proportional to optical forces from
556 non-thermal origin. Other quantities of potential interest, like optical torques
557 and efficiency factors, could also be derived based on the formalism here de-
558 veloped, the latter requiring a more extended discussion (see Sec. 3.13 of
559 Ref. [28]).

560 We expect that this work will serve its purpose in providing a framework
561 for investigating polychromatic light scattering from spherical particles. It
562 now incorporates arbitrary-shaped pulses and multichromatic light gener-
563 ated, for instance, from RGB laser sources, whose total field can be ideally
564 described as a discrete superposition of three monochromatic fields of dif-
565 ferent frequencies. Also, natural extensions of it for multilayered spheres
566 or spheres with inclusions, cylindrical and ellipsoidal particles can also be
567 devised. EBCM, which is a T-matrix method (as other approaches such
568 as GLMT beyond the “strict sense”), can also benefit from the formalism
569 here presented. Finally, we expect that the polychromatic approach of this
570 work will also eventually be incorporated into the analysis of thermal forces
571 like photophoretic forces, now extended for polychromatic light scattering
572 problems.

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