

Regular Article - Theoretical Physics

α + core structure described with an additional interaction in the nuclear matter saturation region

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Received: 13 January 2023 / Accepted: 23 March 2023 / Published online: 12 April 2023

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Communicated by Chong Qi

Abstract In a phenomenological approach, the α + core structure is investigated in the 20 Ne, 44 Ti, 94 Mo, 104 Te, and 212 Po nuclei through the local potential model using a double-folding nuclear potential with effective nucleon-nucleon interaction of M3Y + $c_{\rm sat}\delta(s)$ type, where the term $c_{\rm sat}\delta(s)$ acts only between the saturation regions of α -cluster and core. Properties such as energy levels, α -widths, B(E2) transition rates, and half-lives are calculated, and good level of agreement with experimental data is obtained in general. It is shown the inclusion of the term $c_{\rm sat}\delta(s)$ is determinant for a better description of the experimental energy levels compared to the ground state bands produced with the simple M3Y interaction. The properties calculated for 104 Te reinforce the superallowed α -decay feature for this nucleus, as indicated in previous studies.

1 Introduction

The cluster structure of the atomic nucleus has become one of the main topics in nuclear physics. Its applications are seen in the studies on nuclear structure, spectroscopic properties, nuclear reactions, α or exotic nuclear decay, and astrophysical phenomena. The cluster models are developed with different configurations, ranging from the binary cluster + core structure to molecular structures based on α -particles or α -particles + nucleons. Cluster + core binary systems can be investigated by more direct approaches, such as the local potential model (LPM); on the other hand, the cluster structures can be analyzed by microscopic approaches such as the orthogonality condition model (OCM), the antisymmetrized molecular dynamics (AMD), the resonating group method (RGM), the quantum Monte Carlo method (QMC), and others, mainly in light nuclei [1]. Different experiments have

been carried out to corroborate the theoretical predictions, as shown in recent publications (e.g., [2–5]).

In the last decades, there was an effort by different authors to describe the α + core structure through a systematics applicable to a set of nuclei, mainly nuclei with α -clustering above double-shell closures (e.g., Refs. [6–14]). In general, these works yield satisfactory results for energy spectra, electromagnetic transition rates, α -widths, half-lives, and rms charge radii; however, for some nuclear potential forms, the description of energy levels is unsatisfactory, requiring the use of a variable intensity parameter (in most cases, dependent on the angular momentum L) to obtain energy bands closer to the experimental levels. This feature is observed in the energy bands produced by the double-folding nuclear potentials, mainly in the heavier nuclei. Double-folding nuclear potentials are known to be successful in reproducing α -nucleus elastic scattering cross sections in a wide range of masses and energies, as shown in several works using the DDM3Y interaction [15–19]; however, the double-folding potentials produce energy bands with strong rotational feature, which is incompatible with experimental spectra in many cases; in addition, for heavier nuclei, double-folding potentials can generate very compressed energy bands compared to the experimental levels, making it necessary to use the L-dependent intensity parameter.

The present authors introduced the use of the (1+Gaussian) \times (W.S. + W.S.³) nuclear potential to investigate the α + core structure in 46,54 Cr [20]. Subsequently, the (1 + Gaussian) \times (W.S. + W.S.³) shape was successfully applied to describe the properties of the set $\{^{20}$ Ne, 44 Ti, 94 Mo, 104 Te, 212 Po} [8] and several even-even nuclei of the $22 \le Z \le 42$ region [10]. The recent work of Jia et al. [21] shows that the (1 + Gaussian) \times (W.S. + W.S.³) potential shape can be successfully applied to heavy nuclei other than 212 Po, including the analysis of different cluster structures above 208 Pb. Such a potential has the (1 + Gaussian) factor which favors the correct reproduction of the 0^+ ground state (g.s.) and

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74 Page 2 of 14 Eur. Phys. J. A (2023) 59:74

double-folding with M3Y + $c_{sat}\delta(s)$ (1 + Gaussian)x(WS + WS³)

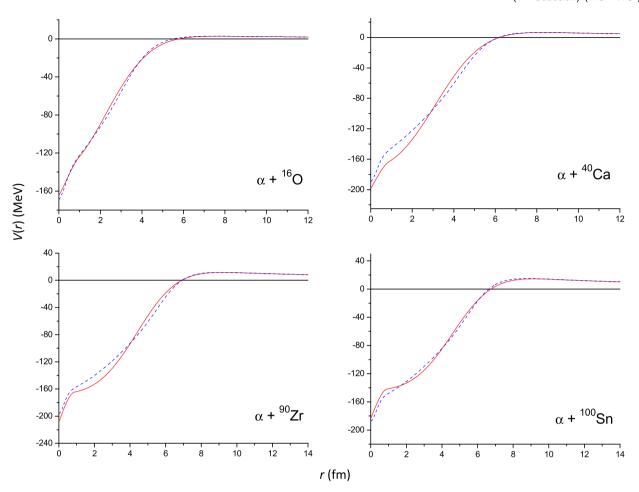


Fig. 1 Potential V(r) applied to the $\alpha+^{16}$ O, $\alpha+^{40}$ Ca, $\alpha+^{90}$ Zr, and $\alpha+^{100}$ Sn systems, where the nuclear term is the double-folding potential with M3Y $+ c_{\text{sat}}\delta(s)$ interaction (solid red line). For each α + core

system, the respective potential composed by the nuclear term of $(1 + Gaussian) \times (W.S. + W.S.^3)$ shape is shown (dashed blue line)

a better description of the $0^+ \to 2^+$ spacing compared to the original W.S. + W.S.³ potential. The (1 + Gaussian) factor produces a depth increase on the W.S. + W.S.³ potential in a region restricted to $r \lesssim 1$ fm (see Fig. 1), acting more strongly on the $0^+ \to 2^+$ spacing; with the inclusion of the centrifugal potential (L > 0), the (1 + Gaussian) factor acts more weakly on the levels above 2^+ . So, the question arises whether the effect produced by the (1 + Gaussian) factor could be described in terms of the effective nucleon-nucleon (NN) interaction, since the double-folding potentials used in previous works do not present such a feature. Furthermore, the reproduction of the (1 + Gaussian) effect on the double-folding potential can provide a more in-depth interpretation of the $(1 + \text{Gaussian}) \times (\text{W.S.} + \text{W.S.}^3)$ potential beyond its properties observed phenomenologically.

The (1 + Gaussian) factor range ($r \lesssim 1$ fm) corresponds to the α -core separation range in which the central regions

of α and core overlap. In the central region of the nucleus (cluster or core), it is known that the nuclear matter density approaches or even exceeds the so-called *saturation density*, $\rho_{\rm sat} \approx 0.17~{\rm fm}^{-3}$. So, it is coherent to associate the (1 + Gaussian) factor range region with the overlap of α and core saturation regions. It should be considered that the superposition of cluster and core mass distributions is an intrinsic property of the double-folding model, as discussed in previous studies on α + nucleus collisions and heavy ion collisions [22,23]. The assumption of overlap of cluster and core saturation regions can lead to a discussion on the Pauli blocking effect at small internuclear separations [24]; however, we restrict this work to a phenomenological approach focused on the systematic description of experimental data.

With such a premise, this work proposes to find a modification in the effective NN interaction which produces an effect similar to the (1 + Gaussian) factor on the nuclear double-



Eur. Phys. J. A (2023) 59:74 Page 3 of 14 **74**

folding potential. In principle, this proposal could be applied to any effective NN interaction form used in previous studies; starting from a simple option, the well-known M3Y-Reid interaction is chosen for this investigation. The M3Y-Reid effective NN interaction is based on the G-matrix elements of the Reid NN potential, and was originally developed as the real part of the optical potential for the description of heavyion scattering [25]. Differently from such use, the present work aims to develop a modified NN interaction for the satisfactory reproduction of energy bands without the use of L-dependence, B(E2) transition rates, α -widths and α -decay half-lives. Therefore, it is natural that the α + core potential proposed in this work presents limitations if applied in the description of α scattering data.

The article is developed as follows: in Sect. 2, a detailed discussion is made on the α + core potential adopted for the set { 20 Ne, 44 Ti, 94 Mo, 104 Te, 212 Po}, the proposed NN effective interaction, the matter density functions for α and core, the quantum number $G_{\rm g.s.}$ applied to the g.s. band of each nucleus, and the similarity between the $(1 + {\rm Gaussian}) \times ({\rm W.S.} + {\rm W.S.}^3)$ potential and the double-folding potential with the proposed modification; in Sect. 3, the results obtained for energy levels, B(E2) transition rates, α -widths, half-lives, rms intercluster separations, and the reduced α -widths are shown, including a comparison of the 104 Te reduced α -widths and the same quantities obtained for 94 Mo and 212 Po; conclusions are presented in Sect. 4.

2 α-cluster model

2.1 α + core potential

The properties of the nucleus are described in terms of an α + core system from the viewpoint of LPM, without considering internal excitations of α -cluster and core. The α + core interaction is described through the local potential

$$V(r) = V_N(r) + V_C(r) \tag{1}$$

containing the nuclear and Coulomb terms. The Coulomb potential $V_C(r)$ is that of an α -particle interacting with an uniformly charged spherical core of radius R,

$$V_C(r) = \begin{cases} \left(\frac{e^2}{4\pi\,\varepsilon_0}\right) \frac{Z_\alpha \, Z_c}{2R} \left(3 - \frac{r^2}{R^2}\right) & \text{for } r < R \\ \left(\frac{e^2}{4\pi\,\varepsilon_0}\right) \frac{Z_\alpha \, Z_c}{r} & \text{for } r \ge R \end{cases}$$
(2)

where $e^2/(4\pi\varepsilon_0) = 1.44$ MeV · fm, Z_α and Z_c are the charge numbers of α -cluster and core, respectively. For the

parameter R, the same values used successfully in the α + core potential of Ref. [8] are employed, since that potential also includes a Coulomb term based on a uniformly charged sphere and is applied for the same set of nuclei. The nuclear potential $V_N(r)$ is given by the double-folding form

$$V_N(r) = \lambda \iint \rho_{\alpha}(\vec{r}_{\alpha})\rho_c(\vec{r}_c)v_{NN}(\vec{s})d\vec{r}_{\alpha}d\vec{r}_c$$
 (3)

where $\vec{s} = \vec{r} - \vec{r}_c + \vec{r}_\alpha$, λ is the intensity parameter, $v_{NN}(\vec{s})$ is the effective NN interaction, $\rho_\alpha(\vec{r}_\alpha)$ and $\rho_c(\vec{r}_c)$ are the nucleon densities of the α -cluster and core, respectively.

As discussed in Sect. 1, it is required that the nuclear potential $V_N(r)$ be such that it becomes more strongly attractive when the saturation regions of α -cluster and core overlap, which occurs in a region restricted to $r \lesssim 1$ fm. Therefore, in the systematics adopted here, we assume that both α and core have a central mass region where $\rho \geq 0.17$ fm⁻³. The proposed shape for $V_N(r)$ can be produced by a suitable modification in the effective NN interaction $v_{NN}(s)$. In this work, a variation of the traditional M3Y-Reid interaction is employed:

$$v_{NN}(s) = V_1 \frac{\exp(-\mu_1 s)}{\mu_1 s} + V_2 \frac{\exp(-\mu_2 s)}{\mu_2 s} + \hat{J}_{00} \delta(s) + c_{\text{sat}} \delta(s).$$
(4)

The first three terms of Eq. (4) are those present in the original form of M3Y interaction, while the term $c_{\text{sat}}\delta(s)$ acts as an attractive interaction *only between the saturation regions of* α -cluster and core. This condition is expressed by:

$$c_{\text{sat}} = 0 \text{ for } \rho_{\alpha}(r_{\alpha}) < 0.17 \text{ fm}^{-3} \text{ or } \rho_{c}(r_{c}) < 0.17 \text{ fm}^{-3};$$

 $c_{\text{sat}} < 0 \text{ for } \rho_{\alpha}(r_{\alpha}) \ge 0.17 \text{ fm}^{-3} \text{ and } \rho_{c}(r_{c}) \ge 0.17 \text{ fm}^{-3}.$

The values for the parameters V_1 , V_2 , μ_1 , and μ_2 are those applied to the M3Y-Reid interaction: $V_1=7999$ MeV, $V_2=2134$ MeV, $\mu_1=4$ fm $^{-1}$, and $\mu_2=2.5$ fm $^{-1}$. In the original M3Y form, \hat{J}_{00} is a factor smoothly dependent on the bombarding energy E_α of the α -particle in the scattering process [25]; however, several energy levels calculated for 20 Ne, 44 Ti, and 94 Mo are below the α + core threshold, making such a dependence on E_α inappropriate in this context. Therefore, we adopted a fixed value $\hat{J}_{00}=-276$ MeV, which corresponds to $E_\alpha=0$ in the original M3Y-Reid form. The parameters λ and $c_{\rm sat}$ are fitted to precisely reproduce the 0^+ and 4^+ levels of the 20 Ne, 44 Ti, 94 Mo, and 212 Po g.s. bands. The values of R, λ and $c_{\rm sat}$ applied to the g.s. bands are shown in Table 1; in addition, this table shows values of the ratio r_0 Coul. $= R/A_L^{1/3}$, where A_T is the mass number of the

 $[\]overline{}^{1}$ Details on the values of λ , $c_{\rm sat}$, and R adopted for 104 Te are discussed in Sect. 3.



74 Page 4 of 14 Eur. Phys. J. A (2023) 59:74

Table 1 Values of the parameters λ , $c_{\rm sat}$, and R, the ratio $r_{0\,\rm Coul.}=R/A_T^{1/3}$, and the quantum number $G_{\rm g.s.}$ applied to the g.s. bands of $^{20}{\rm Ne}, ^{44}{\rm Ti}, ^{94}{\rm Mo}, ^{104}{\rm Te}, {\rm and} ^{212}{\rm Po}$

Nucleus	$G_{ m g.s.}$	λ	c _{sat} (MeV)	R (fm)	r _{0Coul.} (fm)
²⁰ Ne	8	0.88720	-2020	3.272	1.205
⁴⁴ Ti	12	0.79450	-3200	4.551	1.289
⁹⁴ Mo	16	0.71560	-5650	5.783	1.272
¹⁰⁴ Te	16	0.65310	-4870	5.713	1.215
²¹² Po	24	0.79005	-5370	7.018	1.177

total nucleus. Note the small variation of $r_{0\,\text{Coul.}}$ for the set of nuclei, in the range of approx. 1.2-1.3 fm, which is compatible with the relation $R=1.224\,A_T^{1/3}$ (fm) obtained in Ref. [8]. For the negative parity bands, λ is slightly modified for the correct reproduction of the experimental 1^- levels of ^{20}Ne ($E_x=5.7877\,\text{MeV}$) and ^{44}Ti ($E_x=6.220\,\text{MeV}$), and 5 level of ^{94}Mo ($E_x=2.61057\,\text{MeV}$), while the values of R and c_{sat} are those in Table 1 (details in Sect. 3).

In Table 1, it is seen the parameter λ varies in the range of 0.6–0.9, being below the range of 1.1–1.4 applied in similar calculations with the double-folding nuclear potential with DDM3Y interaction [16,18,19,26]. The use of M3Y interaction generates an unnormalized nuclear potential deeper than that obtained with the DDM3Y interaction, as discussed by Kobos et al. [27], and this behavior extends to the M3Y + $c_{\text{sat}}\delta(s)$ interaction; therefore, there is the need for a smaller intensity parameter (λ < 1.0) to reproduce the energy bands of the nuclei considered.

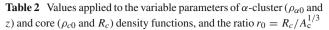
The calculation of the α + core potential depends on determining the density functions $\rho_{\alpha}(r_{\alpha})$ and $\rho_{c}(r_{c})$; moreover, it is necessary to define the saturation region $r_{\alpha(c)} \leq R_{\rm sat}$, where $\rho(r_{\alpha(c)}) \geq 0.17~{\rm fm}^{-3}$, for the density functions of α and core. The radius $R_{\rm sat}$ is determined from the parameter σ of the (1 + Gaussian) \times (W.S. + W.S.³) nuclear potential, as detailed in Sect. 2.2. For the α -cluster density, the 3-parameter Fermi function

$$\rho_{\alpha}(r_{\alpha}) = \rho_{\alpha 0} \left(1 + \frac{w r_{\alpha}^{2}}{c^{2}} \right) / \left[1 + \exp\left(\frac{r_{\alpha} - c}{z}\right) \right]. \tag{5}$$

is used, as proposed in Ref. [28] for the α -particle charge distribution. The original parameter values c=1.01 fm and w=0.445 are applied, while $\rho_{\alpha 0}$ and z are fitted to normalize the total mass to $A_{\alpha}=4$ and reproduce the saturation radius $R_{\rm sat}$ in the α -cluster mass distribution. For the core density, the Woods–Saxon type function is applied:

$$\rho_c(r_c) = \rho_{c0} / \left[1 + \exp\left(\frac{r_c - R_c}{a}\right) \right]$$
 (6)

For the diffuseness parameter, the typical value a=0.55 fm is used, while ρ_{c0} and R_c are fitted to normalize the core mass to A_c nucleons and reproduce the saturation radius $R_{\rm sat}$



System	$\rho_{\alpha 0} (\mathrm{fm}^{-3})$	z (fm)	$\rho_{c0}~(\mathrm{fm}^{-3})$	R_c (fm)	r ₀ (fm)
α +16O	0.19170	0.3677	0.17503	2.4400	0.968
α + 40 Ca	0.19046	0.3690	0.17059	3.5657	1.043
α + 90 Zr	0.19014	0.3693	0.17006	4.8198	1.076
α + 100 Sn	0.19082	0.3686	0.17004	5.0062	1.079
α + 208 Pb	0.19094	0.3685	0.17000	6.4850	1.095

in the core mass distribution. The values of $\rho_{\alpha 0}$, z, ρ_{c0} , and R_c applied to the α + core systems are shown in Table 2; in addition, this table shows values of the ratio $r_0 = R_c/A_c^{1/3}$. There is a small variation of r_0 for the set of nuclei, in the range of approx. 1.0–1.1 fm, which is compatible with the relation $R_c = r_0 A_c^{1/3}$ (fm) expected for the core matter distribution.

2.2 Saturation radius R_{sat}

The saturation radius $R_{\rm sat}$ is estimated from the σ parameter of the (1 + Gaussian) \times (W.S. + W.S.³) nuclear potential. The factor (1 + Gaussian) has the form $1 + \lambda \exp(-r^2/\sigma^2)$, where $\lambda = 0.14$ is one of the fixed parameters of the (1 + Gaussian) \times (W.S. + W.S.³) potential; this factor approaches a constant value = 1 from $r \approx 3\sigma/\sqrt{2}$. From the viewpoint of the double-folding model, the radius $r \approx 3\sigma/\sqrt{2}$ can be interpreted as the cluster-core separation in which the α -cluster and core saturation regions begin to overlap, so that the $c_{\rm sat}\delta(s)$ term of the NN interaction becomes relevant in $r \lesssim 3\sigma/\sqrt{2}$. Such an interpretation is expressed by

$$R_{\operatorname{sat}\alpha} + R_{\operatorname{sat}c} = \frac{3\sigma}{\sqrt{2}},\tag{7}$$

where $R_{\text{sat }\alpha}$ and $R_{\text{sat }c}$ are the saturation radii of α -cluster and core, respectively. Assuming that $R_{\text{sat }\alpha}$ and $R_{\text{sat }c}$ are similar in range, we consider both as a single parameter R_{sat} , resulting in

$$R_{\text{sat}} = \frac{3\sigma}{2\sqrt{2}}. (8)$$

Using the estimated $R_{\rm sat}$ values for the $\alpha+^{16}{\rm O}$, $\alpha+^{40}{\rm Ca}$, $\alpha+^{90}{\rm Zr}$, $\alpha+^{100}{\rm Sn}$, and $\alpha+^{208}{\rm Pb}$ systems, the respective density functions $\rho_{\alpha}(r_{\alpha})$ and $\rho_{c}(r_{c})$ are determined. Figure 2 shows the comparison of the functions $\rho_{c}(r_{c})$ obtained for $^{16}{\rm O}$, $^{40}{\rm Ca}$, $^{90}{\rm Zr}$, and $^{208}{\rm Pb}$ with the respective charge distributions deduced from elastic electron scattering [29], normalized by the A/Z factor. The proximity of the functions $\rho_{c}(r_{c})$ and the respective normalized charge distributions is clearly noted, mainly in the nuclear surface region.



Eur. Phys. J. A (2023) 59:74 Page 5 of 14 **74**

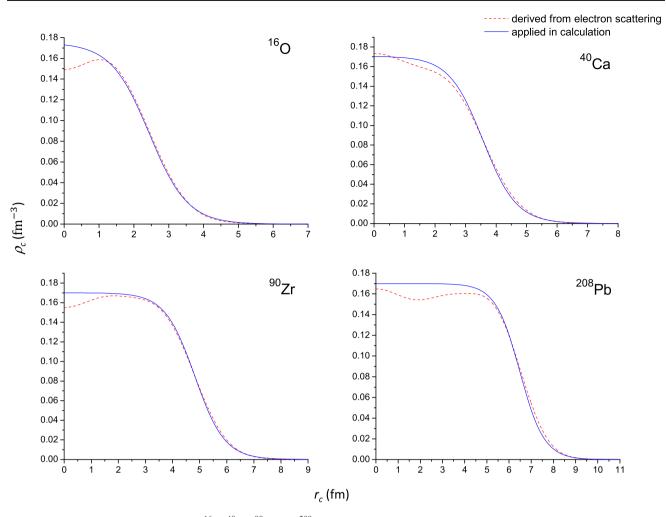


Fig. 2 Nuclear densities $\rho_c(r_c)$ applied to 16 O, 40 Ca, 90 Zr, and 208 Pb in the double-folding calculations (solid blue lines). For comparison, the respective charge distributions normalized by the A/Z factor are

shown (dashed red lines). The charge distributions are deduced from elastic electron scattering and Fourier–Bessel analysis [29]

2.3 Quantum number $G_{g.s.}$

The α -cluster nucleons must lie in shell-model orbitals outside the core. This restriction is defined through the global quantum number G = 2N + L, where N is the number of internal nodes in the radial wave function and L is the orbital angular momentum; thus, the restriction is expressed by $G \geq G_{g.s.}$, where $G_{g.s.}$ corresponds to the g.s. band. The numbers $G_{g.s.} = 8$ and 12 are applied to the g.s. bands of 20 Ne and 44 Ti, respectively, and $G_{g.s.} = 16$ is applied to 94 Mo and 104 Te, according to the Wildermuth condition [30]. $G_{g.s.} = 24$ is applied to the g.s. band of ²¹²Po, as this number provides a satisfactory reproduction of the experimental energy levels from 0⁺ to 8⁺ using the double-folding potential, including the compression trend of the energy spacing from the 4^+ level. The number $G_{g.s.} = 24$ has been suitably used in previous work as [31,32] to describe the α -decay half-life of 212 Po, using other α + core potential shapes.

2.4 Comparison of double-folding and $(1 + Gaussian) \times (W.S. + W.S.^3)$ potentials

Using the methodology proposed in Sects. 2.1, 2.2, and 2.3, the α + core potential was determined for analysis of 20 Ne, 44 Ti, 94 Mo, 104 Te, and 212 Po. A comparison of the α + core potentials obtained with the double-folding and $(1 + \text{Gaussian}) \times (\text{W.S.} + \text{W.S.}^3)$ nuclear terms is shown in Fig.1 for 20 Ne, 44 Ti, 94 Mo, and 104 Te. Note that the two potential forms are similar for the analyzed nuclei, including the $r \leq 1$ fm region where both become more intensely attractive. Therefore, it is shown that the effect of the (1 + Gaussian) factor can be reproduced in the double-folding potential with the proposed modification in the effective NN interaction. For 212 Po, such a comparison is not presented since there is a difference between the $G_{g.s.}$ number applied in this work with the double-folding potential $(G_{g.s.} = 24)$ and that applied with



74 Page 6 of 14 Eur. Phys. J. A (2023) 59:74

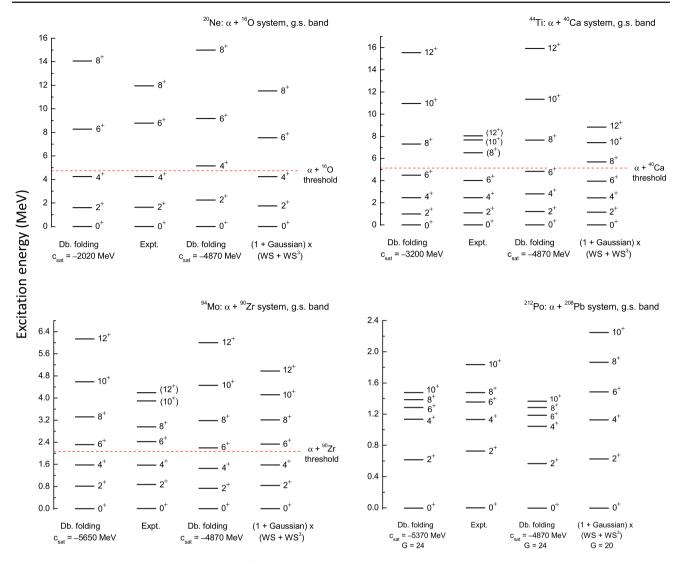


Fig. 3 Ground state bands calculated for the $\alpha+^{16}$ O ($G_{g.s.}=8$), $\alpha+^{40}$ Ca ($G_{g.s.}=12$), $\alpha+^{90}$ Zr ($G_{g.s.}=16$), and $\alpha+^{208}$ Pb ($G_{g.s.}=24$) systems using the double-folding nuclear potential with M3Y + $c_{sat}\delta(s)$ interaction. For each nucleus, the g.s. band is calculated firstly with c_{sat} fitted for the correct reproduction of the 0^+ and 4^+ states, and secondly,

with $c_{\text{sat}} = -4870$ MeV. The fixed value of the intensity parameter (λ) applied for each α + core system is shown in Table 1. In addition, the g.s. bands calculated with the (1 + Gaussian) × (W.S. + W.S.³) nuclear potential [8] are shown for comparison

the $(1 + \text{Gaussian}) \times (\text{W.S.} + \text{W.S.}^3)$ potential in Ref. [8] $(G_{\text{g.s.}} = 20)$.

3 Results

The g.s. bands of 20 Ne, 44 Ti, 94 Mo, and 212 Po are calculated with the α + core potential defined in Sect. 2 and the parameter values in Table 1. The resolution of Schrödinger radial equation for the α + core relative motion allows to determine the energy levels of the system and respective radial wave functions. Figure 3 shows the comparison of calculated and experimental energy levels. It is shown the α + core poten-

tial provides a good description of experimental levels from 0^+ to 8^+ . For L>8, the description of the experimental levels is rougher, as the double-folding potential produces an energy band with rotational behavior at higher spins for 20 Ne, 44 Ti, and 94 Mo. However, it should be noted that the satisfactory results from 0^+ to 8^+ were obtained without using an L-dependent intensity parameter. As shown in Fig. 3, the double-folding and $(1 + \text{Gaussian}) \times (\text{W.S.} + \text{W.S.}^3)$ potentials produce similar bands for 20 Ne, 44 Ti, and 94 Mo from 0^+ to 8^+ , and a more visible difference at L>8, where the $(1 + \text{Gaussian}) \times (\text{W.S.} + \text{W.S.}^3)$ potential gives a better description of the experimental energy levels for 44 Ti and 94 Mo. In the case of 212 Po, the double-folding potential pro-



Eur. Phys. J. A (2023) 59:74 Page 7 of 14 **74**

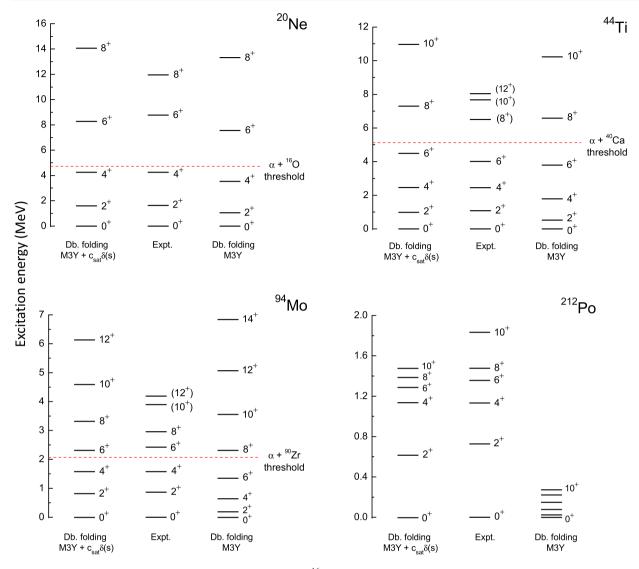


Fig. 4 Comparison of the ground state bands calculated for the $\alpha+^{16}$ O ($G_{\rm g.s.}=8$), $\alpha+^{40}$ Ca ($G_{\rm g.s.}=12$), $\alpha+^{90}$ Zr ($G_{\rm g.s.}=16$), and $\alpha+^{208}$ Pb ($G_{\rm g.s.}=24$) systems using the double-folding nuclear potential with the

effective NN interactions: M3Y and M3Y $+ c_{\text{sat}} \delta(s)$. For each nucleus, c_{sat} is fitted for the correct reproduction of the 0^+ and 4^+ states

duces an energy band compressed at L > 4, which favors a better reproduction of the experimental band compared to the $(1 + \text{Gaussian}) \times (\text{W.S.} + \text{W.S.}^3)$ potential.

In a next calculation, a single value of the parameter $c_{\rm sat}$ is determined for the 20 Ne, 44 Ti, 94 Mo, and 212 Po nuclei. For this, $c_{\rm sat}$ is varied to minimize the quantity

$$S^2 = \sum (E_{\text{expt}} - E_{\text{calc}})^2 \tag{9}$$

for the energy levels from 0^+ to 8^+ of the 20 Ne, 44 Ti, 94 Mo, and 212 Po g.s. bands. Using this procedure, the value $c_{\rm sat} = -4870$ MeV is obtained. In Fig. 3, it is seen the bands calculated with the fixed value $c_{\rm sat} = -4870$ MeV are still satisfactory in describing the levels from 0^+ to 8^+ , showing it is possible to describe the four nuclei through an effective NN interaction with fixed parameters.

It is interesting to compare the g.s. bands calculated with and without the $c_{\rm sat}\delta(s)$ term of the effective NN interaction, to verify its level of importance in reproducing the experimental levels. Figure 4 shows the g.s. bands of 20 Ne, 44 Ti, 94 Mo, and 212 Po calculated with the effective NN interactions: M3Y-Reid (applying $\hat{J}_{00}=-276$ MeV) and the $v_{NN}(s)$ interaction of Eq. (4). In the case of 20 Ne and 44 Ti, it is seen that the contribution of the $c_{\rm sat}\delta(s)$ term is relevant for a better reproduction of the experimental levels from 0^+ to 4^+ . However, in the case of 94 Mo, the contribution of the $c_{\rm sat}\delta(s)$ term is more significant, providing a better reproduction of the experimental levels from 0^+ to 8^+ and canceling the rotational behavior produced with the M3Y interaction from 0^+ to 8^+ . In the case of 212 Po, the contribution of the $c_{\rm sat}\delta(s)$ term is evident. For this nucleus, the g.s. band pro-



74 Page 8 of 14 Eur. Phys. J. A (2023) 59:74

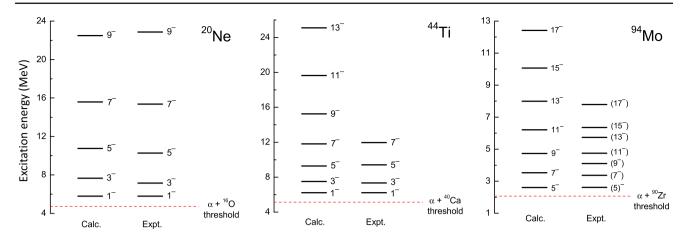


Fig. 5 Negative parity bands calculated for the α +¹⁶O (G=9), α +⁴⁰Ca (G=13), and α +⁹⁰Zr (G=17) systems through the double-folding nuclear potential with M3Y + $c_{\rm sat}\delta(s)$ interaction, in compar-

ison with experimental energy levels [33,34]. The intensity parameter λ is slightly modified for the correct reproduction of the 1⁻ levels of ²⁰Ne and ⁴⁴Ti, and (5)⁻ of ⁹⁴Mo (the λ values are discussed in the text)

duced with the simple M3Y interaction is very compressed and incompatible with the experimental band; e.g., the calculated 10^+ level is at $E_x=0.273$ MeV, which is far below the respective experimental energy. With the inclusion of the $c_{\rm sat}\delta(s)$ term in the effective NN interaction, the calculated band provides a good description of the experimental levels from 0^+ to 8^+ , also reproducing the compression trend of the experimental band from the 4^+ level. As the compression of the 212 Po calculated band remains above 8^+ , only the experimental 10^+ level is more roughly reproduced.

The model also allows the calculation of excited bands of the α + core systems. Figure 5 shows the calculated negative parity bands for 20 Ne (G = 9), 44 Ti (G = 13), and 94 Mo $(G = 17)^2$ compared to experimental energies. For calculating these bands, the values of the intensity parameter of $V_N(r)$ are: $\lambda = 0.9025$, 0.8050, and 0.7703 for ²⁰Ne, ⁴⁴Ti, and 94 Mo, respectively. The relative variation of λ from the g.s. band to the negative parity band is very small for ²⁰Ne (1.72 %) and ⁴⁴Ti (1.32 %), showing the α + core potential is weakly dependent on the band number G for these two nuclei. In the case of 94 Mo, the relative variation of λ is somewhat greater (7.64 %), and the experimental negative parity band is more compressed compared to the calculated band; however, the calculation for ⁹⁴Mo should be interpreted as exploratory, since the spins and parities of the considered experimental levels are uncertain.

The g.s. band calculated for 104 Te is shown in Fig. 6, case (b). In this calculation, the parameter λ is fitted to reproduce

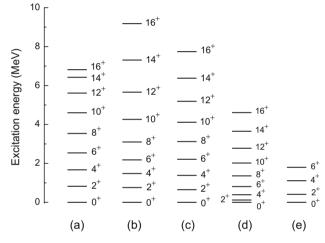


Fig. 6 Ground state band calculated for the $\alpha + ^{100}{\rm Sn}$ system $(G_{\rm g.s.} = 16)$ in the cases: (a) using the nuclear potential of $(1+{\rm Gaussian}) \times ({\rm W.S.} + {\rm W.S.}^3)$ shape [8]; (b) using the double-folding nuclear potential with M3Y + $c_{\rm sat}\delta(s)$ effective interaction (this work); (c) calculation of P. Mohr [35] using the double-folding nuclear potential with DDM3Y interaction and L-dependent intensity parameter; (d) calculation of Bai and Ren [36] using the double-folding nuclear potential with M3Y interaction and fixed intensity parameter; (e) calculation of Jiang et al. [37] using the shell-model without isospin symmetry (SM1)

the energy $E(0^+)=5.22$ MeV used in Ref. [8] to describe the ground state of 104 Te. In addition, $c_{\rm sat}=-4870$ MeV is adopted, since this value proved to be satisfactory for the set $\{^{20}$ Ne, 44 Ti, 94 Mo, 212 Po}. As in the other nuclei, the value of R used in Ref. [8] for the Coulomb potential is applied (R=5.713 fm for 104 Te). In Fig. 6, the 104 Te g.s. band calculated in this work is compared with previous calculations in the excitation energy scale. Note that the energy levels calculated through the double-folding nuclear potential with M3Y + $c_{\rm sat}\delta(s)$ interaction are close to the respective J^+ levels obtained with the (1 + Gaussian) × (W.S. + W.S. 3) nuclear potential from 0^+ to 12^+ , indicating the similarity



² The ⁹⁴Mo experimental negative parity band presents levels from $(5)^-$ ($E_x = 2.611$ MeV) to (17^-) ($E_x = 7.782$ MeV) connected by γ -transitions [33]; however, the excited state $(5)^-$ undergoes a γ -transition to the 4⁺ state ($E_x = 1.574$ MeV) of the G = 16 g.s. band. As the excited state $(5)^-$ does not undergo a transition of $J \to J - 2$ type, the G = 17 band is assumed to be incomplete and the calculation of the corresponding theoretical band is made from $J^\pi = 5^-$.

Eur. Phys. J. A (2023) 59:74 Page 9 of 14 **74**

Table 3 Calculated α_0 -widths ($\Gamma_{\alpha_0}^{\text{calc.}}$) for the resonant states of the ground state band (G=8) and negative parity band (G=9) of ²⁰Ne in comparison with experimental data ($\Gamma_{\alpha_0}^{\text{expt.}}$) [43]. An α preformation factor $P_{\alpha}=1$ is considered

J^{π}	G	E_x (MeV)	$\Gamma_{\alpha_0}^{\text{calc.}}$ (keV)	$\Gamma_{\alpha_0}^{\text{expt.}}$ (keV)
6+	8	8.7776	1.08	0.11 ± 0.02
8+	8	11.951	0.74	0.035 ± 0.010
1-	9	5.7877	0.061	0.028 ± 0.003
3-	9	7.1563	18.5	8.2 ± 0.3
5-	9	10.262	302	145 ± 40
7 ^{-a}	9	13.692	240	158 ± 15
7-	9	15.366	890	78 ± 7
9-a	9	17.430	37.2	53 ± 6
9-	9	22.87	1.36×10^{3}	$(0.225 \pm 0.040) \times 10^3$

^a Energy levels not shown in Fig. 5

of the two potential shapes. Furthermore, there is a clear proximity (mainly from 0^+ to 12^+) between the band calculated with the M3Y + $c_{sat}\delta(s)$ interaction and the band obtained by Mohr [35] through the double-folding potential with DDM3Y interaction and L-dependent λ parameter. The band calculated by Bai and Ren [36] is more compressed and has strongly rotational behavior, which is produced by the double-folding potential with simple M3Y interaction. Also, it is seen the calculation of the present work generates energy spacings $2^+ \rightarrow 4^+$ and $4^+ \rightarrow 6^+$ similar to the shellmodel calculation without isospin symmetry (SM1) by Jiang et al. [37].³ The evaluation of the calculated energy bands is limited by the absence of experimental data related to the excited states of ¹⁰⁴Te; nevertheless, the comparison of the different 104 Te theoretical bands indicates the $c_{\text{sat}}\delta(s)$ term added to the effective NN interaction produces an effect on energy levels similar to that generated by the L-dependent intensity parameter applied to the double-folding potential (except at the highest spin levels).

The α -decay width (Γ_{α}) of 104 Te is calculated by the semiclassical approximation of Ref. [38], using an α preformation factor $P_{\alpha}=1$; such an approximation is derived from the two-potential approach [38–40], which is often applied in α decay studies. The α -decay width obtained for the ground state at $E(0^+)=5.22$ MeV is $\Gamma_{\alpha}=1.0188\times 10^{-13}$ MeV, and the corresponding α -decay half-life, given by $T_{1/2,\alpha}=\hbar \ln 2/\Gamma_{\alpha}$, is ≈ 4.478 ns. This $T_{1/2,\alpha}$ value is compatible with the experimental measure of Auranen et al. [41] ($T_{1/2}^{\rm exp}<18$ ns). By varying the parameter λ , it is verified that the energy range 5.044 MeV $< E(0^+) < 5.3$ MeV provides $T_{1/2,\alpha}$ values compatible with the experimental measure, and simultaneously, is compatible with the experimental measure $O_{\alpha}^{\rm exp}=5.1(2)$ MeV. In agreement with the calculation in Ref. [8], the production of such results with $P_{\alpha} = 1$ corroborates the indication of superallowed α -decay of ¹⁰⁴Te by Auranen et al.

The α_0 -decay widths calculated for the ²⁰Ne resonant states are shown in Table 3 in comparison with experimental values. The widths are calculated by the semiclassical approximation of Ref. [38], except for the 9⁻ level $(E_x = 22.87 \text{ MeV})$, where the α_0 -width is calculated by $\Gamma_{\alpha} = 2/(d\delta_L/dE)|_{E=E_r}$, where δ_L is the scattering phase shift for a given L, and E_R is the resonance energy. In this calculation, the parameter λ is smoothly modified so that the experimental energy of the J^{π} state is reproduced precisely. The Γ_{α_0} values calculated for the 6⁺ and 8⁺ levels of the g.s. band have an order of magnitude above that observed for the respective experimental values; these results repeat a discrepancy found in some previous calculations on ²⁰Ne [6,11,12] with different α + core potentials. Concerning the G=9 band, the calculated Γ_{α_0} values correctly reproduce the order of magnitude of the experimental α_0 -widths of 1⁻, 3^{-} , 5^{-} , 7^{-} ($E_x = 13.692$ MeV), and 9^{-} ($E_x = 17.430$ MeV) states. Such results corroborate the statement of previous works [6,42] in which the 7^- ($E_x = 13.692$ MeV) and 9^- ($E_x = 17.430$ MeV) states are better candidates for members of the G = 9 band of the $\alpha + {}^{16}$ O system; however, it is not excluded that both 7⁻ and both 9⁻ experimental levels in Table 3 can be associated with the theoretical 7⁻ and 9^- states of the G=9 band.

Table 4 shows the γ -decay widths (Γ_{γ}) , α -decay widths (Γ_{α}) , α -branching ratios (b_{α}) , and half-lives $(T_{1/2})$ calculated for the states from 0^+ to 8^+ of the 212 Po g.s. band. The Γ_{γ} values are calculated through eq. (9) of Ref. [26], and the Γ_{α} values are calculated by the semi-classical approximation of Ref. [38]. To calculate Γ_{γ} , it is necessary to previously determine the B(E2) transition rates for the 212 Po g.s. band; the B(E2) rates calculated for 212 Po are shown in Table 6. Also, the value of Γ_{γ} depends on the internal conversion coefficient α_{IC} (see values in Table 4), which is obtained from the internal conversion coefficient database BrIcc [45]. An α pre-



 $^{^3}$ Ref. [37] presents other shell-model calculations for the 104 Te g.s. band, such as the SD-pair approximation, the SDG1K-pair approximation and the shell-model calculation with isospin symmetry (SM2); however, such calculations produce results similar to the SM1 calculation shown in Fig. 6.

74 Page 10 of 14 Eur. Phys. J. A (2023) 59:74

Table 4 γ -decay widths (Γ_{γ}) , α -decay widths (Γ_{α}) , α -branching ratios (b_{α}) , and half-lives $(T_{1/2})$ calculated for the states from 0^+ to 8^+ of the ²¹²Po g.s. band, in comparison with experimental data ([34], except

where indicated). The values adopted for the internal conversion coefficient (α_{IC}) are obtained from Ref. [45]. An α preformation factor $P_{\alpha} = 3.902 \%$ is applied

J^{π}	E_x (MeV)	α_{IC}	$\Gamma_{\gamma}(J \to J - 2) (\text{MeV})$	Γ_{α} (MeV)	$b_{\alpha}^{\mathrm{calc.}}\left(\%\right)$	$b_{\alpha}^{\mathrm{expt.}}\left(\%\right)$	$T_{1/2}^{\mathrm{calc.}}$	$T_{1/2}^{\text{expt.}}$
0+	0.0000			1.550×10^{-15}	100.000	100	294.3 ns	294.3(8) ns
2^+	0.7273	0.01393	9.406×10^{-11}	4.061×10^{-14}	0.043	0.033	4.8 ps	14.2(18) ps
4+	1.1325	0.0543	7.418×10^{-12}	8.224×10^{-14}	1.097	≈ 0.5	60.8 ps	69(10) ps [2,44] ^a
6^+	1.3555	0.324	4.880×10^{-13}	3.273×10^{-14}	6.285	$\approx 3(1)$	0.88 ns	0.76(21) ns ^b
8+	1.4764	3.248	6.994×10^{-14}	4.226×10^{-15}	5.698	$\approx 3(1)$	6.2 ns	14.6(3) ns

^aThe recent experimental work of Karayonchev et al. [46] presents the measure $T_{1/2}^{\text{expt.}}(4^+) = 70(5)$ ps for ²¹²Po

formation factor $P_{\alpha}=3.902$ % is applied to correctly reproduce the experimental half-life of 212 Po in the 0^+ ground state ($T_{1/2}^{\rm expt.}=294.3(8)$ ns); this P_{α} value is close to the one used by Hoyler, Mohr, and Staudt [26] ($P_{\alpha}=3.5$ %) for the same nucleus. The α -branching ratio is obtained by $b_{\alpha}=\Gamma_{\alpha}/(\Gamma_{\alpha}+\Gamma_{\gamma})$.

Analyzing the results in Table 4, it is noted the calculated α -branching ratios reproduce the order of magnitude of the respective experimental data; in the case of $b_{\alpha}(6^{+})$ and $b_{\alpha}(8^{+})$, such a statement takes into account the proximity between the calculated values and the experimental uncertainty range of ± 1 %. The half-lives $T_{1/2}$ calculated for the 2^+ , 4^+ , 6^+ , and 8^+ levels also reproduce the order of magnitude of the respective experimental values and, in the cases of $T_{1/2}(4^+)$ and $T_{1/2}(6^+)$, the calculated values are in complete agreement with the respective experimental data. It should be mentioned that the recent experimental work of von Tresckow et al. [2] presents the measure $T_{1/2}^{\text{expt.}}(^{212}\text{Po}; 6^+) = 1.15(19)$ ns, and the recent experimental work of Karayonchev et al. [46] presents the measures $T_{1/2}^{\text{expt.}}(^{212}\text{Po};4^+) = 70(5)$ ps and $T_{1/2}^{\text{expt.}}(^{212}\text{Po},6^+) =$ 1.14(10) ns, which are still in good level of agreement with the respective calculated values.

Table 5 shows the rms intercluster separations $(\langle R^2 \rangle^{1/2})$ and reduced α -widths (γ_{α}^2) calculated for the levels from 0^+ to 8^+ of the g.s. bands of the studied nuclei. These properties are calculated according to the procedure detailed in Subsection IV.B of our previous work in Ref. [13]. For obtaining these properties, the parameter λ is slightly modified (see λ values in Table 5) so that the radial wave function u(r) of each J^{π} state is calculated at the respective experimental energy, except to 104 Te where the intensity parameter has the fixed value $\lambda=0.65310$ mentioned in Table 1; for 20 Ne, 44 Ti, 94 Mo, and 212 Po, such a procedure aims to determine the nuclear properties as $\langle R^2 \rangle^{1/2}$, γ_{α}^2 , B(E2), etc. in correspondence with the experimental energy levels. Analyzing the rms intercluster separations, it is noted that the antistretching effect, which is observed with other α + core potential

shapes (e.g., [11,13,16,18,47]), also occurs with the potential proposed in this work; thus, the inclusion of the $c_{\text{sat}}\delta(s)$ term in the effective NN interaction does not change this general property. The reduced α -widths, together with the rms intercluster separations, indicate the first levels of each g.s. band have a stronger α -cluster character.

Table 6 shows the calculated B(E2) transition rates between the levels from 0⁺ to 8⁺ of the ²⁰Ne, ⁴⁴Ti, ⁹⁴Mo, and ²¹²Po g.s. bands. The calculated B(E2) values are compared with the respective experimental data and B(E2) results from previous studies [6,7,9,11,12] in which the α + core structure is systematically analyzed for a set of nuclei. The results of the present work provide a good general description of the experimental B(E2) data, since the order of magnitude of most of the experimental data is reproduced correctly, and concerning the absolute values, there are cases of complete agreement between calculated and experimental values for the transitions: $4^+ \rightarrow 2^+$ and $6^+ \rightarrow 4^+$ of ²⁰Ne; $6^+ \to 4^+ \text{ of } ^{44}\text{Ti}; 4^+ \to 2^+ \text{ and } 6^+ \to 4^+ \text{ of } ^{212}\text{Po.}^4 \text{ Also,}$ in the $8^+ \rightarrow 6^+$ transition of ²⁰Ne, the calculated B(E2)value is very close to the uncertainty range of the respective experimental measure. Such results are rewarding, taking into account they were obtained without introducing effective charges. In Table 6, it is seen the B(E2) values of the present work are considerably higher than those obtained in Refs. [6,7,9,11,12] with different α + core potentials; consequently, the B(E2) values of this work are somewhat closer to the experimental data in most transitions. Thus, the B(E2)results favor the α + core interpretation for 20 Ne, 44 Ti, 94 Mo, and ²¹²Po by the methodology proposed here.

Table 7 presents the reduced α -widths calculated for the g.s. band of ^{212}Po with $G_{\text{g.s.}}=22$. The alterna-



^bThe recent experimental works of von Tresckow et al. [2] and Karayonchev et al. [46] present the measures $T_{1/2}^{\text{expt.}}(6^+) = 1.15(19)$ ns and 1.14(10) ns for ²¹²Po, respectively

⁴ The recent experimental works of von Tresckow et al. [2] and Karayonchev et al. [46] present the measures $B(E2; 6_1^+ \rightarrow 4_1^+) = 8.7(15)$ W.u. and 9.0(+9-7) W.u. for ²¹²Po, respectively, which favors the theoretical predictions of Refs. [7,9,12]. However, the 2020 ²¹²Po experimental data compilation [34,50] presents the value $B(E2; 6_1^+ \rightarrow 4_1^+) = 13.2(+49-29)$ W.u. shown in Table 6.

Eur. Phys. J. A (2023) 59:74 Page 11 of 14 **74**

Table 5 Rms intercluster separations $(\langle R^2 \rangle^{1/2})$ and reduced α-widths (γ_α^2) calculated for the 0^+ to 8^+ levels of the 20 Ne $(G_{g.s.}=8)$, 44 Ti $(G_{g.s.}=12)$, 94 Mo $(G_{g.s.}=16)$, 104 Te $(G_{g.s.}=16)$, and 212 Po $(G_{g.s.}=24)$ ground state bands. The strength parameter λ fitted for the precise reproduction of each experimental energy is shown, except to 104 Te where the strength parameter has the fixed value $\lambda=0.65310$

Nucleus	J^{π}	λ	$\langle R^2 \rangle^{1/2}$ (fm)	$\gamma_{\alpha}^{2} (\text{keV})$
²⁰ Ne	0+	0.8872	4.018	39.26
	2+	0.8863	4.050	43.73
	4+	0.8872	4.009	38.65
	6+	0.8765	3.971	34.19
	8+	0.9286	3.500	6.97
⁴⁴ Ti	0^{+}	0.7945	4.461	4.32
	2+	0.7934	4.468	4.63
	4 ⁺	0.7945	4.427	3.78
	6+	0.7999	4.316	2.16
	8+	0.8034	4.180	1.02
⁹⁴ Mo	0^{+}	0.71560	5.158	0.68
	2+	0.71514	5.168	0.75
	4 ⁺	0.71560	5.137	0.60
	6+	0.71473	5.067	0.37
	8+	0.71854	4.949	0.15
¹⁰⁴ Te	0+	0.65310	5.469	2.00
	2+	0.65310	5.480	2.19
	4+	0.65310	5.447	1.79
	6+	0.65310	5.351	1.05
	8+	0.65310	5.227	0.49
²¹² Po	0+	0.79005	6.559	0.63
	2+	0.78920	6.577	0.72
	4+	0.79007	6.556	0.61
	6+	0.78953	6.494	0.40
	8+	0.78940	6.405	0.21

tive use of $G_{\rm g.s.}=22$ for $^{212}{\rm Po}$ aims at a comparison with the $\gamma_{\alpha}^2(^{212}{\rm Po})$ results obtained with $G_{\rm g.s.}=24$ and shown in Table 5. Using the $\gamma_{\alpha}^2(^{104}{\rm Te})$ values, the ratios $\gamma_{\alpha}^2(^{104}{\rm Te})/\gamma_{\alpha}^2(^{94}{\rm Mo})$, $\gamma_{\alpha}^2(^{104}{\rm Te})/\gamma_{\alpha}^2(^{212}{\rm Po})$ with $G_{\rm g.s.}(^{212}{\rm Po})=24$ were calculated (see Table 7). The $\gamma_{\alpha}^2(^{104}{\rm Te})/\gamma_{\alpha}^2(^{94}{\rm Mo})$ values of this work are considerably higher than those shown in Ref. [8], where the reduced α -widths were obtained with the (1 + Gaussian) × (W.S. + W.S.³) potential; however, it should be taken into account that small variations on the length or height of the Coulomb barrier can produce strong variations in the γ_{α}^2 values. The use of $G_{\rm g.s.}=22$ produces $\gamma_{\alpha}^2(^{212}{\rm Po})$ values lower than in the case of $G_{\rm g.s.}(^{212}{\rm Po})=24$, resulting in higher $\gamma_{\alpha}^2(^{104}{\rm Te})/\gamma_{\alpha}^2(^{212}{\rm Po})$ ratios for $G_{\rm g.s.}(^{212}{\rm Po})=22$. The $\gamma_{\alpha}^2(^{104}{\rm Te})/\gamma_{\alpha}^2(^{212}{\rm Po})$ ratios in Table 7 are lower than those obtained in Ref. [8]; however, the calculation of Ref. [8] applies $G_{\rm g.s.}(^{212}{\rm Po})=20$, which naturally contributes to higher values of $\gamma_{\alpha}^2(^{104}{\rm Te})/\gamma_{\alpha}^2(^{212}{\rm Po})$. Even with such quan-

titative differences, the calculations of the present work reinforce the interpretation of Ref. [8] in which 104 Te has a substantially higher degree of α -clustering compared to 94 Mo and 212 Po. Such results also corroborate the analysis of Auranen et al. [41] based on experimental 104 Te α -decay data.

4 Summary and conclusions

This work shows a systematic study of the α + core structure in 20 Ne, 44 Ti, 94 Mo, 104 Te, and 212 Po nuclei using a double-folding nuclear potential with effective NN interaction of M3Y + $c_{\rm sat}\delta(s)$ type, where the term $c_{\rm sat}\delta(s)$ acts only between the α and core saturation regions. Using this nuclear potential plus the Coulomb term, a good description of the experimental g.s. bands of 20 Ne, 44 Ti, 94 Mo, and 212 Po is obtained, mainly from $J^{\pi}=0^+$ to 8^+ . With a slight variation of the λ intensity parameter, the experimental negative parity bands of 20 Ne and 44 Ti are well reproduced. It is shown the additional term $c_{\rm sat}\delta(s)$ of the effective NN interaction is determinant for a better reproduction of the experimental



74 Page 12 of 14 Eur. Phys. J. A (2023) 59:74

Table 6 B(E2) transition rates calculated for the levels from 0^+ to 8^+ of the 20 Ne $(G_{g.s.} = 8)$, 44 Ti $(G_{g.s.} = 12)$, 94 Mo $(G_{g.s.} = 16)$, and 212 Po $(G_{g.s.} = 24)$ ground state bands, without the use of effective of the state of th

tive charges. The B(E2) values of the present work are compared with results from previous studies on the α + core structure. Experimental data are from Ref. [34], except where indicated

Nucleus	J^{π}	$B(E2; J \rightarrow J - 2)$ (W.u.)							
		Expt.	This work	Buck et al. [6]	Ibrahim et al. [11]	Bai and Ren [12]	Wang et al. [7]	Ni and Ren [9]	
²⁰ Ne	2+	20.3(10)	16.646	14.6	14.3	13.0	13.3	12.2	
	4+	22(2)	22.686	18.6	18.5	17.1	17.7	15.8	
	6^+	20(3)	21.883	14.9	15.2	15.2	15.2	13.6	
	8+	9.0(13)	10.834	7.4	7.9	7.0	8.0	6.2	
⁴⁴ Ti	2+	22.2(+ 22 -18) [48]	11.291	11.0	10.6	9.9	10.1	9.3	
	4+	30(+4-3)[49]	15.423	14.7	14.1	13.4	13.7	12.6	
	6^+	15(+3-2)[49]	14.662	14.1	13.2	12.7	13.4	11.8	
	8+		12.229	11.5	10.4	10.5	11.2	9.6	
⁹⁴ Mo	2+	16.0(4)	8.075	7.8	7.5			6.8	
	4+	26(4)	11.184	10.7	10.2			9.3	
	6^+		11.232		10.2			9.3	
	8+	0.0049(8)	9.915		9.2			8.2	
²¹² Po	2+	2.57(+ 38 -29)	7.521			6.3	6.7	5.8 ^a	
	4+	9.4(13) [2] ^b	10.632			8.8	9.5	8.2^{a}	
	6^+	$13.2(+49-29)^c$	11.035			9.1	10.0	8.5^{a}	
	8+	4.56(12)	10.519			8.7	9.8	8.1^{a}	

^a B(E2) rates calculated with G=22

Table 7 Reduced α-widths calculated for the g.s. band of 212 Po ($G_{g.s.} = 22$) from $J^{\pi} = 0^+$ to 8^+ , and ratios $\gamma_{\alpha}^2(^{104}$ Te)/ $\gamma_{\alpha}^2(^{94}$ Mo), $\gamma_{\alpha}^2(^{104}$ Te)/ $\gamma_{\alpha}^2(^{212}$ Po) in the case $G_{g.s.}(^{212}$ Po) = 22, and $\gamma_{\alpha}^2(^{104}$ Te)/ $\gamma_{\alpha}^2(^{212}$ Po) in the case $G_{g.s.}(^{212}$ Po) = 24

J^{π}	$\gamma_{\alpha}^{2}(^{212}\text{Po}; G_{\text{g.s.}} = 22) \text{ (keV)}$	$\gamma_{\alpha}^2(^{104}\text{Te})/\gamma_{\alpha}^2(^{94}\text{Mo})$	$\gamma_{\alpha}^{2}(^{104}\text{Te})/\gamma_{\alpha}^{2}(^{212}\text{Po}) [G_{\text{g.s.}}(^{212}\text{Po}) = 22]$	$\gamma_{\alpha}^{2}(^{104}\text{Te})/\gamma_{\alpha}^{2}(^{212}\text{Po})$ $[G_{\text{g.s.}}(^{212}\text{Po}) = 24]$
0+	0.393	2.935	5.092	3.196
2^+	0.451	2.906	4.849	3.050
4+	0.379	2.990	4.712	2.918
6^+	0.241	2.863	4.369	2.619
8+	0.119	3.261	4.092	2.319

levels, and produces an effect similar to the (1 + Gaussian) factor of the $(1 + Gaussian) \times (W.S. + W.S.^3)$ nuclear potential employed in previous works [8,10,20,21].

The proposed potential V(r) is used to analyze the resonant energy levels of $^{20}\mathrm{Ne}$ and $^{212}\mathrm{Po}$, and the α -decay half-life of $^{104}\mathrm{Te}$. Concerning $^{20}\mathrm{Ne}$, the Γ_{α_0} widths calculated for the G=9 negative parity band reproduce the order of magnitude of most experimental data. In the case of $^{212}\mathrm{Po}$, the α -branching ratios and half-lives calculated for the g.s. band reproduce the order of magnitude of the experimental data with a preformation factor $P_{\alpha}\approx 3.9$ %, and there is complete agreement with the experimental half-lives of the 4^+ and 6^+ states. With respect to $^{104}\mathrm{Te}$, the model shows agree-

ment with the experimental measurements of Q_{α} and $T_{1/2,\alpha}$ in the energy range 5.044 MeV $< E(0^+) < 5.3$ MeV using an α preformation factor $P_{\alpha}=1$; this result corroborates the indication of superallowed α -decay of 104 Te in previous studies.

The B(E2) transition rates calculated for the g.s. bands of $^{20}\mathrm{Ne}$, $^{44}\mathrm{Ti}$, $^{94}\mathrm{Mo}$, and $^{212}\mathrm{Po}$ present a good level of agreement with the experimental data, as the order of magnitude of most experimental B(E2) rates is reproduced correctly, and there is complete agreement with the experimental data in some transitions without the use of effective charges. The B(E2) results of the present work are somewhat closer to the



^bThe recent experimental work of Karayonchev et al. [46] presents the measure $B(E2; 4_1^+ \to 2_1^+) = 9.3(+7 - 6)$ W.u. for ²¹²Po

^cThe recent experimental works of von Tresckow et al. [2] and Karayonchev et al. [46] present the measures $B(E2; 6_1^+ \to 4_1^+) = 8.7(15)$ W.u. and 9.0(+9-7) W.u. for 212 Po, respectively

Eur. Phys. J. A (2023) 59:74 Page 13 of 14 **74**

experimental rates than other systematic calculations on the α + core structure in previous studies.

The calculated reduced α -widths are used to determine the ratios $\gamma_{\alpha}^2(^{104}\text{Te})/\gamma_{\alpha}^2(^{94}\text{Mo})$, $\gamma_{\alpha}^2(^{104}\text{Te})/\gamma_{\alpha}^2(^{212}\text{Po})$ with $G_{\text{g.s.}}(^{212}\text{Po})=22$, and $\gamma_{\alpha}^2(^{104}\text{Te})/\gamma_{\alpha}^2(^{212}\text{Po})$ with $G_{\text{g.s.}}(^{212}\text{Po})=24$; the obtained values reinforce the interpretation of Souza et al. [8] in which ^{104}Te has an α -clustering degree substantially higher than ^{94}Mo and ^{212}Po , and corroborate the analysis by Auranen et al. [41] based on experimental ^{104}Te α -decay data.

In conclusion, the double-folding nuclear potential with M3Y + $c_{\text{sat}}\delta(s)$ effective NN interaction provides a good general description of energy levels, B(E2) rates, α -widths and α -decay half-lives through a systematics applied to the set $\{^{20}\text{Ne},\,^{44}\text{Ti},\,^{94}\text{Mo},\,^{104}\text{Te},\,^{212}\text{Po}\}$. Additionally, this calculation allows reproducing the effect of the (1 + Gaussian) factor of the $(1 + \text{Gaussian}) \times (\text{W.S.} + \text{W.S.}^3)$ nuclear potential from the viewpoint of the double-folding model. The successful use of the $c_{\text{sat}}\delta(s)$ term in the effective NN interaction suggests the importance of a further discussion on this phenomenological approach in terms of subnucleonic interactions.

Acknowledgements The authors thank the HPC resources provided by Information Technology Superintendence (HPC-STI) of University of São Paulo. Support from Instituto Nacional de Ciência e Tecnologia – Física Nuclear e Aplicações (INCT-FNA) is acknowledged.

Data Availability Statement This manuscript has no associated data or the data will not be deposited. [Authors' comment: As a theoretical physics publication, there is no experimental data to be shared.]

References

- M. Freer, H. Horiuchi, Y. Kanada-Enyo, D. Lee, U.-G. Meißner, Rev. Mod. Phys. 90, 035004 (2018)
- M. von Tresckow, M. Rudigier, T.M. Shneidman, T. Kröll, M. Boromiza, C. Clisu et al., Phys. Lett. B 821, 136624 (2021)
- 3. Y.-Z. Wang, S. Zhang, Y.-G. Ma, Phys. Lett. B 831, 137198 (2022)
- S. Bailey, T. Kokalova, M. Freer, C. Wheldon, R. Smith, J. Walshe, N. Soić, L. Prepolec, V. Tokić, F.M. Marqués, L. Achouri, F. Delaunay, M. Parlog, Q. Deshayes, B. Fernández-Dominguez, B. Jacquot, Eur. Phys. J. A 57, 108 (2021)
- P. Dasgupta, G.-L. Ma, R. Chatterjee, L. Yan, S. Zhang, Y.-G. Ma, Eur. Phys. J. A 57, 134 (2021)
- 6. B. Buck, A.C. Merchant, S.M. Perez, Phys. Rev. C 51, 559 (1995)
- 7. S.M. Wang, J.C. Pei, F.R. Xu, Phys. Rev. C 87, 014311 (2013)
- 8. M.A. Souza, H. Miyake, T. Borello-Lewin, C.A. da Rocha, C. Frajuca, Phys. Lett. B **793**, 8 (2019)
- 9. D. Ni, Z. Ren, Phys. Rev. C 83, 014310 (2011)
- 10. M.A. Souza, H. Miyake, Phys. Rev. C 104, 064301 (2021)
- T.T. Ibrahim, A.C. Merchant, S.M. Perez, B. Buck, Phys. Rev. C 99, 064332 (2019)
- 12. D. Bai, Z. Ren, Phys. Rev. C 103, 044316 (2021)
- 13. M.A. Souza, H. Miyake, Phys. Rev. C 91, 034320 (2015)
- 14. P. Mohr, Open Nucl. Part. Phys. J. 1, 1 (2008)
- A.M. Kobos, B.A. Brown, R. Lindsay, G.R. Satchler, Nucl. Phys. A 425, 205 (1984)

- 16. S. Ohkubo, Phys. Rev. Lett. 74, 2176 (1995)
- P. Mohr, G.G. Kiss, Z. Fülöp, D. Galaviz, G. Gyürky, E. Somorjai, At. Data Nucl. Data Tables 99, 651 (2013)
- 18. P. Mohr, Eur. Phys. J. A 53, 209 (2017)
- U. Atzrott, P. Mohr, H. Abele, C. Hillenmayer, G. Staudt, Phys. Rev. C 53, 1336 (1996)
- 20. M.A. Souza, H. Miyake, Eur. Phys. J. A 53, 146 (2017)
- 21. J. Jia, Y. Qian, Z. Ren, Phys. Rev. C 104, L031301 (2021)
- D.T. Khoa, G.R. Satchler, W. von Oertzen, Phys. Rev. C 56, 954 (1997)
- 23. Y. Sakuragi, Prog. Theor. Exp. Phys. 2016, 06A106 (2016)
- V.B. Soubbotin, W. von Oertzen, X. Viñas, K.A. Gridnev, H.G. Bohlen, Phys. Rev. C 64, 014601 (2001)
- 25. G.R. Satchler, W.G. Love, Phys. Rep. 55, 183 (1979)
- 26. F. Hoyler, P. Mohr, G. Staudt, Phys. Rev. C 50, 2631 (1994)
- A.M. Kobos, B.A. Brown, P.E. Hodgson, G.R. Satchler, A. Budzanowski, Nucl. Phys. A 384, 65 (1982)
- R.F. Frosch, J.S. McCarthy, R.E. Rand, M.R. Yearian, Phys. Rev. 160, 874 (1967)
- Nuclei Charge Density Archive, and references therein. http://discovery.phys.virginia.edu/research/groups/ncd/index.html
- K. Wildermuth, Y.C. Tang, A Unified Theory of the Nucleus (Academic Press, New York, 1977)
- B. Buck, A.C. Merchant, S.M. Perez, J. Phys. G Nucl. Part. Phys. 17, 1223 (1991)
- 32. B. Buck, A.C. Merchant, S.M. Perez, Phys. Rev. C 45, 2247 (1992)
- 33. Y.H. Zhang, M. Hasegawa, W.T. Guo, M.L. Liu, X.H. Zhou, G. de Angelis, T.M. Axiotis, A. Gadea, N. Marginean, Martinez, D.R. Napoli, C. Rusu, Z. Podolyak, C. Ur, D. Bazzacco, F. Brandolini, S. Lunardi, S.M. Lenzi, R. Menegazzo, R. Schwengner, A. Gargano, W. von Oertzen, S. Tazaki, Phys. Rev. C 79, 044316 (2009)
- 34. ENSDF: Evaluated Nuclear Structure Data File, and references therein. https://www.nndc.bnl.gov/ensdf
- 35. P. Mohr, Eur. Phys. J. A 31, 23 (2007)
- 36. D. Bai, Z. Ren, Eur. Phys. J. A 54, 220 (2018)
- H. Jiang, C. Qi, Y. Lei, R. Liotta, R. Wyss, Y.M. Zhao, Phys. Rev. C 88, 044332 (2013)
- 38. S.A. Gurvitz, G. Kalbermann, Phys. Rev. Lett. **59**, 262 (1987)
- Y. Qian, Z. Ren, D. Ni, J. Phys. G Nucl. Part. Phys. 38, 015102 (2011)
- 40. Y. Oian, Z. Ren, D. Ni, Phys. Rev. C 83, 044317 (2011)
- K. Auranen, D. Seweryniak, M. Albers, A.D. Ayangeakaa, S. Bottoni, M.P. Carpenter et al., Phys. Rev. Lett. 121, 182501 (2018)
- B. Buck, J.C. Johnston, A.C. Merchant, S.M. Perez, Phys. Rev. C 52, 1840 (1995)
- D.R. Tilley, C.M. Cheves, J.H. Kelley, S. Raman, H.R. Weller, Nucl. Phys. A 636, 249 (1998)
- XUNDL: Experimental Unevaluated Nuclear Data List Search and Retrieval, and references therein. https://www.nndc.bnl.gov/xundl
- T. Kibédi, T.W. Burrows, M.B. Trzhaskovskaya, P.M. Davidson, C.W. Nestor Jr., Nucl. Instr. Methods Phys. Res. A 589, 202 (2008). http://bricc.anu.edu.au
- V. Karayonchev, G. Rainovski, J. Jolie, A. Blazhev, A. Dewald, A. Esmaylzadeh, C. Fransen, P. John, L. Knafla, D. Kocheva, K. Schomacker, V. Werner, H. Naïdja, Phys. Rev. C 106, 064305 (2022)
- F. Michel, S. Ohkubo, G. Reidemeister, Prog. Theor. Phys. Suppl. 132, 7 (1998)
- K. Arnswald, T. Braunroth, M. Seidlitz, L. Coraggio, P. Reiter, B. Birkenbach et al., Phys. Lett. B 772, 599 (2017)
- K. Arnswald, P. Reiter, A. Blazhev, T. Braunroth, A. Dewald, M. Droste, C. Fransen, A. Goldkuhle, R. Hetzenegger, R. Hirsch, E. Hoemann, L. Kaya, L. Lewandowski, C. Müller-Gatermann, P. Petkov, D. Rosiak, M. Seidlitz, B. Siebeck, A. Vogt, D. Werner, K. Wolf, K.-O. Zell, Phys. Rev. C 102, 054302 (2020)
- 50. K. Auranen, E.A. McCutchan, Nucl. Data Sheets 168, 117 (2020)



74 Page 14 of 14 Eur. Phys. J. A (2023) 59:74

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